

Improvements of Geomagnetic Field Models Made During the First Years of the International Decade

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- Satellite Data during the International Decade
- Examples of Recent Field Models
- Satellite Data vs. Observatory Data
- What can be expected from Swarm?

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Availability of Satellite Data: The Last 40 Years



Geopotential Research

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Present Satellites of the International Decade: Ørsted, CHAMP, SAC-C

Ørsted

First high-precision geomagnetic satellite since Magsat (1979-80) Launched on 23th February 1999 Polar orbit, 650-850 km altitude all local times within 790 days (2.2 years)

• CHAMP

Launched on 15th July 2000 low altitude (400-450 km) all local times within 130 days

• SAC-C Ørsted-2

Copy of Ørsted experiment Launched on 21th November 2000 700 km altitude, fixed local time 10³⁰/22³⁰ (no high-precision vector data due to payload failure)







Availability of Satellite Data during the International Decade





Models of Earth's Magnetic Field

Geophysical

VOLUME IN NUMBER 15

Station for Britsserven & Line

101

Research

Letters

- More and more data
- Improved methods





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Example of Field Models:

• CM4 (Sabaka et al., 2004)



• MF4 (Maus et al., 2006)





• CHAOS (Olsen et al., 2006)

Comprehensive Model CM4: Parameterize fields from all major near- Earth sources

- Simultaneous estimation of the time-space structure of
 - main field (core + crust),
 - ionospheric field,
 - magnetospheric field,
 - induced field,

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field due to currents at satellite height.



Data

- POGO (scalar), 1965-71
- Magsat (vector + scalar), 1979-80
- Ørsted (vector + scalar), 1999-2002.5
- CHAMP (scalar), 2002.5-2002.5
- observatory data, 1960-2001
- about 2.2 mio observations
- about 25,000 model parameters

Sabaka et al., 2004



Examples of CM4 Results

crustal field ...

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... and fit to observed North component

(at 1am on geomagnetic quietest day of each month) for Fürstenfeldbruck and Kakioka





Lithospheric Field Model MF4

- Derived from almost 5 years of CHAMP data
 - scalar dat at all latitudes
 - vector data equatorward of 65° dipole latitude
- Removal of oceanic tidal signal



- Track-by track high-pass filtering, to remove large-scale magnetospheric field
- Line leveling
- "Data super selection": Removal of
 - 30% outliers (non polar latitudes)
 - 95% outliers (polar latitudes)

Spherical harmonic expansion up to degree/order *n*=90 (regularized for *n* > 60)

Maus et al., 2006



Example of MF4 Result: Scalar anomaly at 50 km altitude



Maus et al., 2006



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Maus et al., 2006



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CHAOS -CHAMP, Oersted, SAC-C Magnetic Field Model

- Model of Earth's Magnetic Field based on almost 7 years of satellite data (March 99 – Dec 05)
- Goal of CHAOS "optimal" description of
 - static field, up to n = 40
 - "high-degree" first time derivative (linear secular variation), up to n = 16
 - time-space structure of core field (temporal spline representation), up to *n*=14
- Multi-satellite vector data alignment
 - Scalar data
 - Vector components in m
 - Attitude data
- *in magnetometer frame in star imager frame*

Vector data alignment is done as part of the field modeling effort by co-estimation of alignment parameters instead of serial estimation





CHAOS Model Parameterization

- Model of core and crustal fields:
 - static: $N_{MF} = 50$
 - first time derivative: $N_{SV} = 18$
 - spline parameterisation: N_{spline} = 14:
 (1 yr knot spacing)
- Model of magnetospheric fields
 - separate estimation large-scale magnetospheric field every 12 hours:

4752 coefficients

3113 coefficients

Vector data alignment (Euler angles)

3 x 113 parameters

• 58% of the model is "target 'field" (core and crustal field)





Result: Power Spectrum of Secular Variation

Power Spectrum: mean square magnetic field (at Earth's surface) due to contributions with spherical harmonic degree n (horizontal wavelength λ_n)



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dY/dt at Niemegk/Germany







dY/dt at Niemegk/Germany





Evidence for a Jerk around 2003 in LOD

- Length of Day (LOD) curve
- Occurence of "Wiggles" in d(LOD)/dt are correlated with jerks (Holme & de Viron, 2005)



Time change of LOD suggest occurence of jerk in 1999 and 2003



Distribution of observatories, 1999-2005



- Observatory data (e.g. monthly means) provide an excellent opportunity to study the magnetic field change at a given point in space
- But very uneven spatial distribution of observatories
- Is it possible to determine secular variation from satellite data obtained at or near the grid points of a regular net ?
- Monthly means at "virtual observatories"
 in space, similar to observatory monthly means ?





Determination of secular variation (monthly means) at given location in space from satellite data in vicinity (e.g., 400 km radius) of "target point".



Mandea and Olsen, in review

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Monthly Means at ground and at 400 km altitude: Three observatories, all three components ...



Satellite and observatory data contain well correlated signal that is not described by present field models. Is this signal internal or external in origin?

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Regular grid of virtual observatories at 400 km altitude: *Z*-component



monthly means CHAOS model prediction





Regular grid of virtual observatories at 400 km altitude: dZ/dt



monthly means CHAOS model prediction



External Field Contributions in Observatory Monthly Means

2005

2005

2005

2004

2004

2004

vear

year

year





observatory monthly means

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model predictions (based on satallite data only): only internal field (CHAOS, modified CHAOS) internal and external field field

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What can be expected from Swarm?

Lithospheric Field and Gradient at 400 km altitude







0 nT 20



-20





 B_{ϕ}

East-West Gradient of Magnetic field



-0.02



Improvement of Lithospheric Field Model



POGO and Magsat ... N = 40, resolution: 1000 km

Magnetic field of Earth's crust radial component at 10 km altitude





Improvement of Lithospheric Field Model



... with present satellites Ørsted and CHAMP ... N = 65, resolution: 615 km

Magnetic field of Earth's crust radial component at 10 km altitude





Improvement of Lithospheric Field Model



... and with Swarm N = 150, resolution: 270 km

Magnetic field of Earth's crust radial component at 10 km altitude

