

Swarm and Gravity

Possibilities and Expectations for Gravity Field Recovery

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- **Representation of the Gravity Field and Spaceborne Observations**

- **Methodology for Error Analysis**
 - error propagation using transfer coefficients
 - full simulation based on ...
 - energy integral method

- **Use of Swarm for Gravity Field Recovery**

- **Simulation Results**

Gravity Field and Spaceborne Observations

Representation of the gravity field in terms of **spherical harmonics (sh)**

On the sphere:

$$\begin{aligned}
 V(r, \theta, \lambda) &= \frac{GM}{R} \sum_{l=0}^{L_{\max}} \left(\frac{R}{r}\right)^{l+1} \sum_{m=0}^l \bar{P}_{lm}(\cos \theta) (\bar{C}_{lm} \cos m\lambda + \bar{S}_{lm} \sin m\lambda) \\
 &= \frac{GM}{R} \sum_{l=0}^{L_{\max}} \left(\frac{R}{r}\right)^{l+1} \sum_{m=-l}^l \bar{P}_{lm}(\cos \theta) \bar{K}_{lm} e^{im\lambda}
 \end{aligned}$$

dimensioning upward cont. gravitational potential coefficients

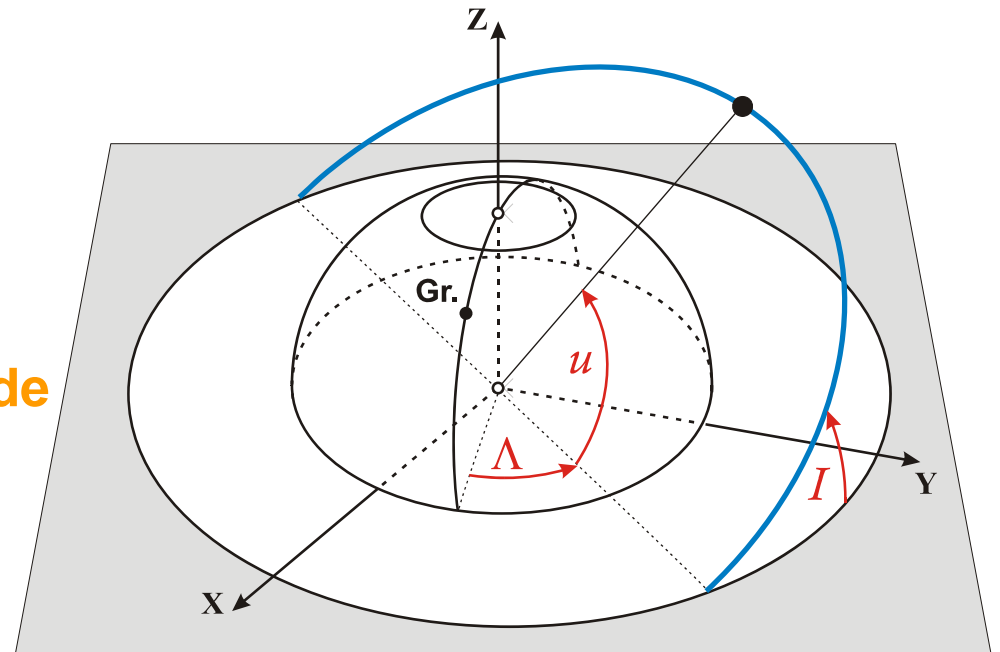
Gravity Field and Spaceborne Observations

Along the orbit:

$$V(r, u, \Lambda) = \frac{GM}{R} \sum_{l=0}^{L_{\max}} \left(\frac{R}{r} \right)^{l+1} \sum_{m=-l}^l \sum_{k=-l}^l \bar{K}_{lm} \bar{F}_{lmk}(I) e^{i(ku+m\Lambda)}$$

orbit coordinates:

- r** - geocentric radius
- u** - argument of latitude
- Λ** - longitude of asc. node
- I** - inclination



Gravity Field and Spaceborne Observations

Representation along the orbit for arbitrary functionals $f^\#$:

$$f^\#(r, u, \Lambda) = \sum_m \sum_k A_{mk}^\# e^{i(ku+m\Lambda)}$$

with lumped coefficients

$$A_{mk}^\# = \sum_l H_{lmk}^\# \bar{K}_{lm}$$

and transfer coefficients

$$H_{lmk}^\#$$

for example:

$$H_{lmk}^V = \frac{GM}{R} \left(\frac{R}{r} \right)^{l+1} \bar{F}_{lmk}(I)$$

where # can be, e.g., gravitational potential

V

gravity gradients

V_{ij}

(GOCE)

orbit perturbations

$\Delta \mathbf{x}$

(CHAMP)

range, range rate

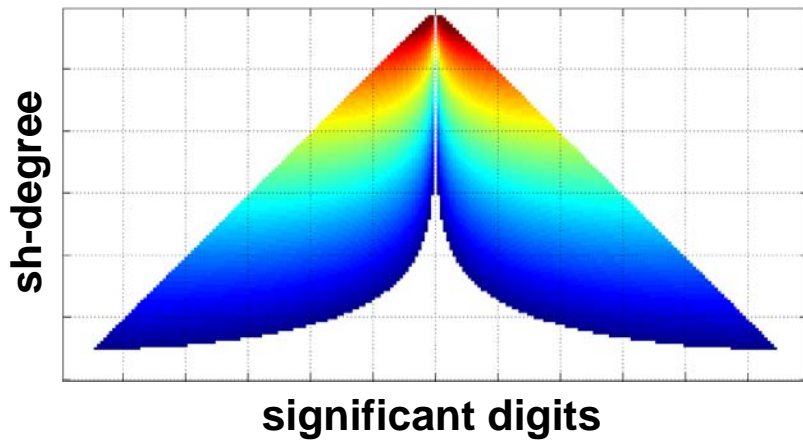
$\rho, \dot{\rho}$

(GRACE)

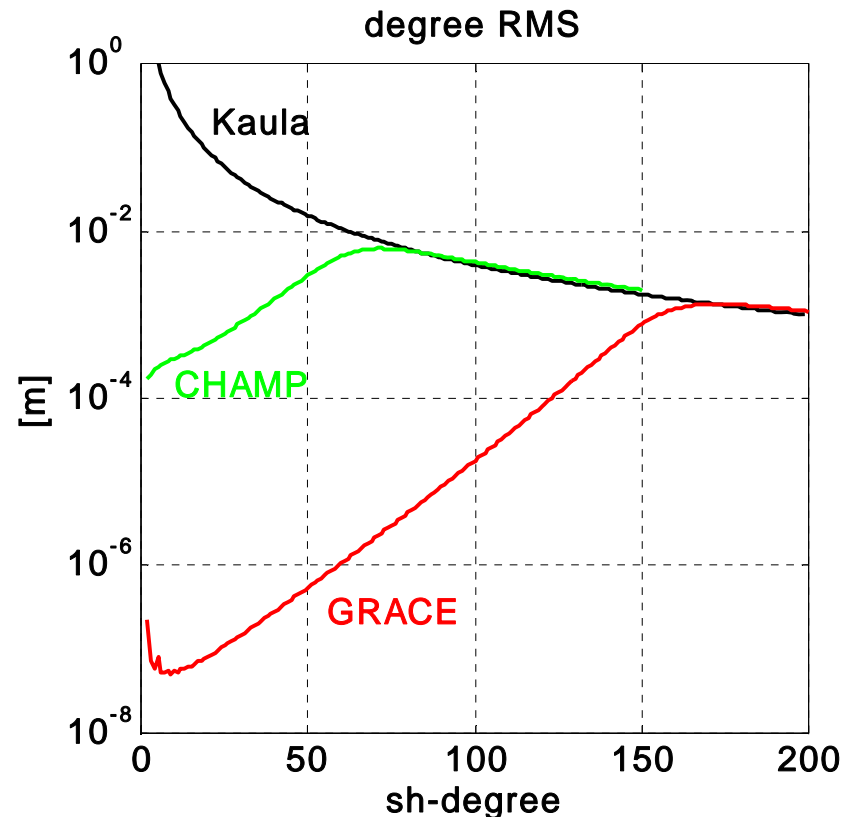
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Methodology – Error Propagation

- based on error models of the observations
- using specific transfer coefficients $H_{lmk}^{\#}$

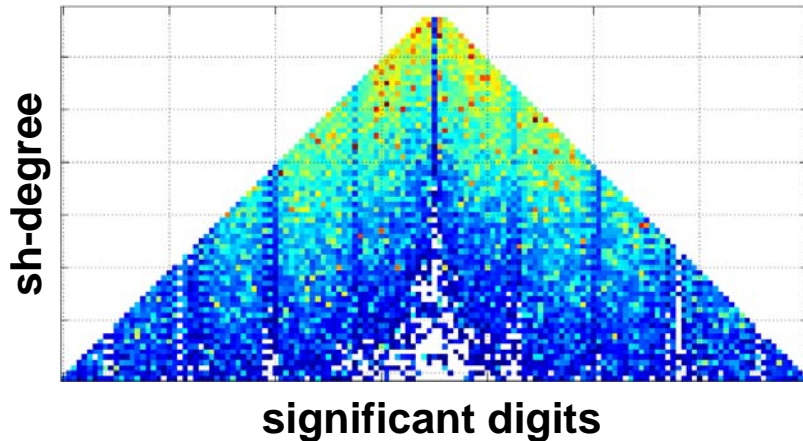


⇒ formal error estimates (variances)

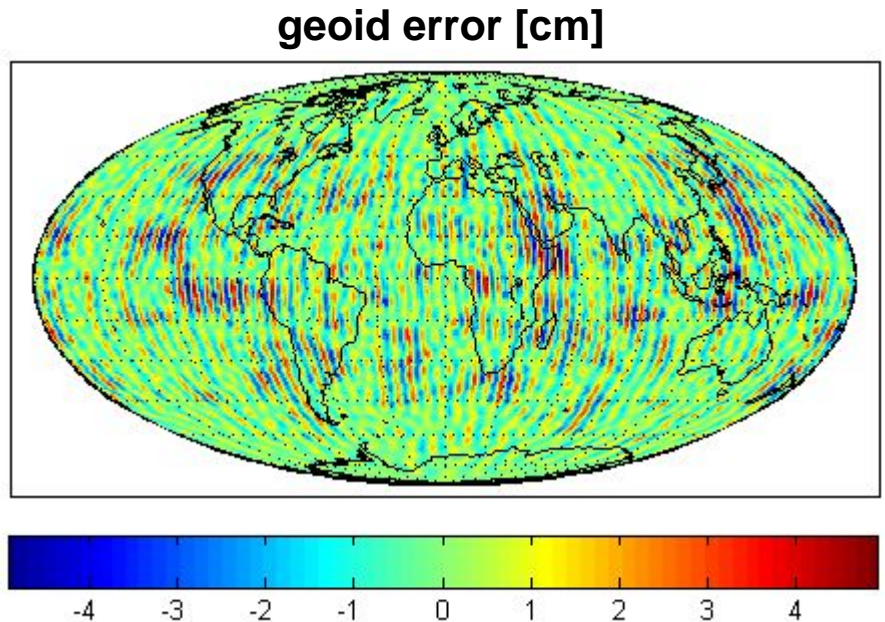


Methodology – Full Simulation

- based on **energy integral** \Rightarrow **pseudo observation** $V, \Delta V$
- using time series of observation errors



\Rightarrow **empirical errors**
(differences to „truth“)



Methodology - Energy Integral (one satellite)

Energy conservation law

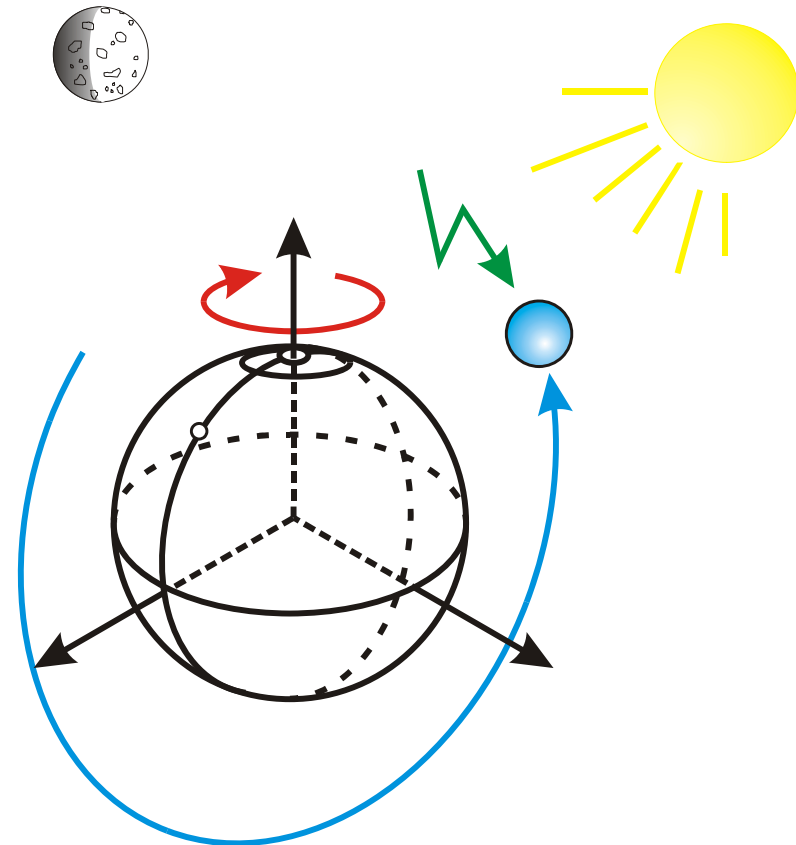
$$E_{kin} + E_{pot} = const.$$

$$\Rightarrow V = \frac{\dot{\mathbf{x}}^2}{2} - Z + \int_x \mathbf{a} \cdot d\mathbf{x} + C$$

centrifugal
potential

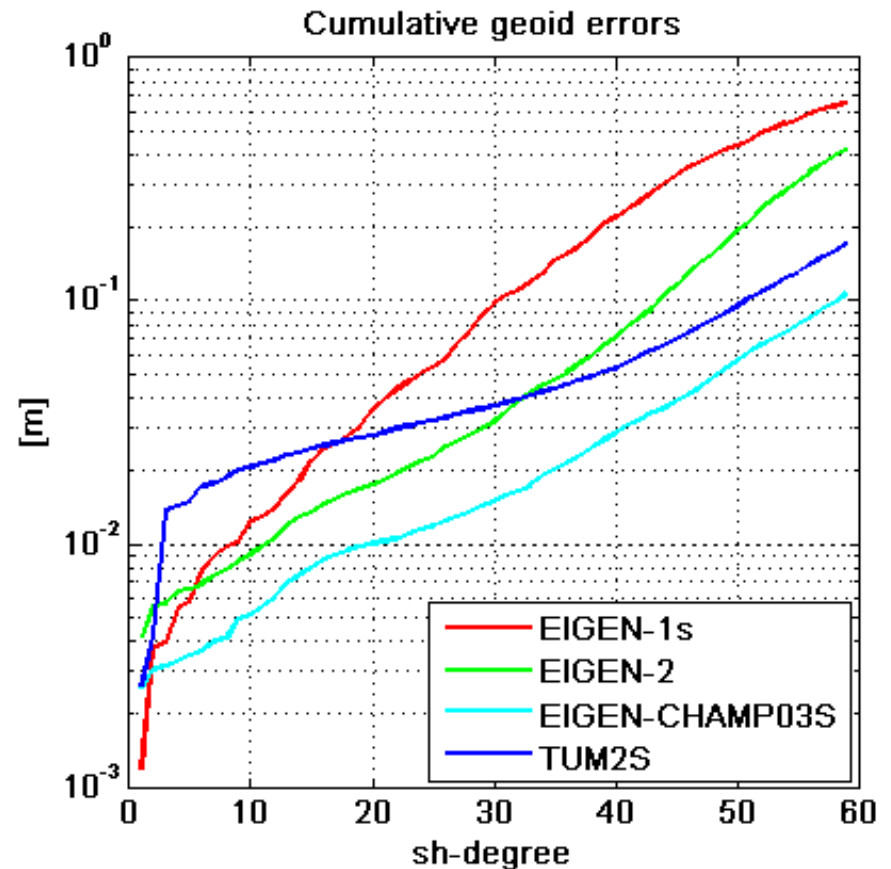
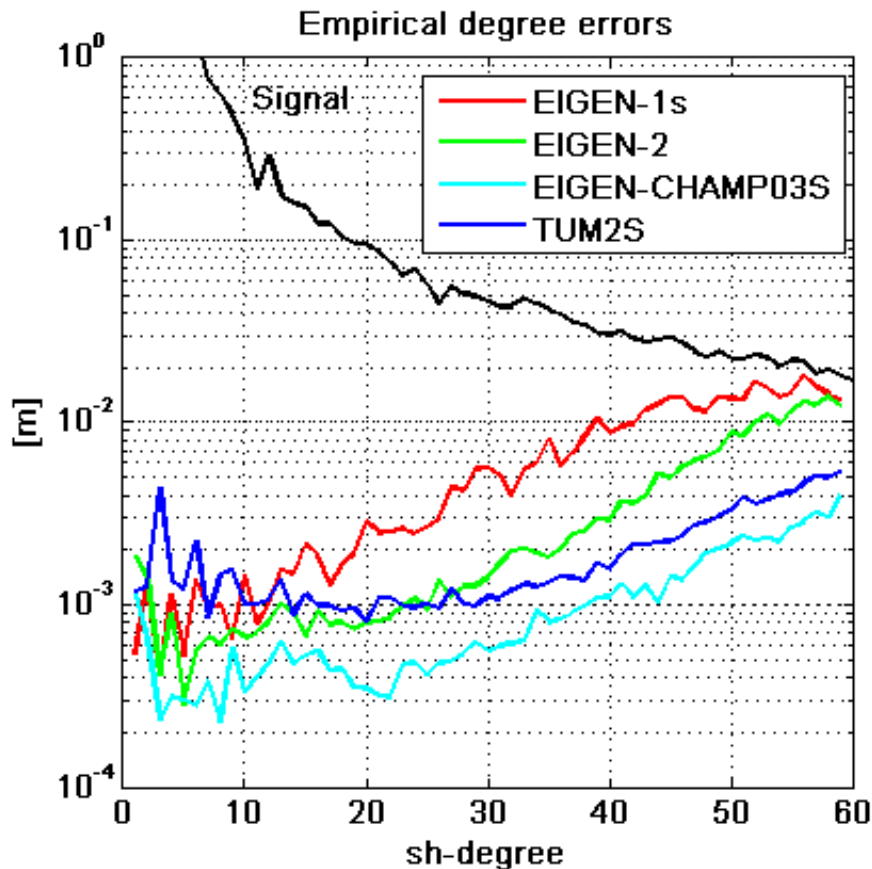
non-conservative
forces

$$V \hat{=} \frac{\dot{\mathbf{x}}^2}{2}$$



Methodology - Energy Integral (proof of concept)

Global gravity models by IAPG using CHAMP data
(empirical errors wrt. EIGEN-GRACE02S)

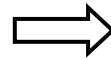


Methodology - Energy Integral (two satellites)

Relative motion between satellites

$$V \hat{=} \frac{\dot{\mathbf{x}}^2}{2}$$

absolute velocity vector



$$\Delta V \hat{=} \frac{\Delta \dot{\mathbf{x}}^2}{2} + \dot{\mathbf{x}}_1 \cdot \Delta \dot{\mathbf{x}}$$

relative velocity vector

For **along-track** separation

GRACE: scalar KBR-observation

$$\Delta V \approx \|\dot{\mathbf{x}}_1\| \dot{\rho}$$

range rate

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Swarm and Gravity - Constellation

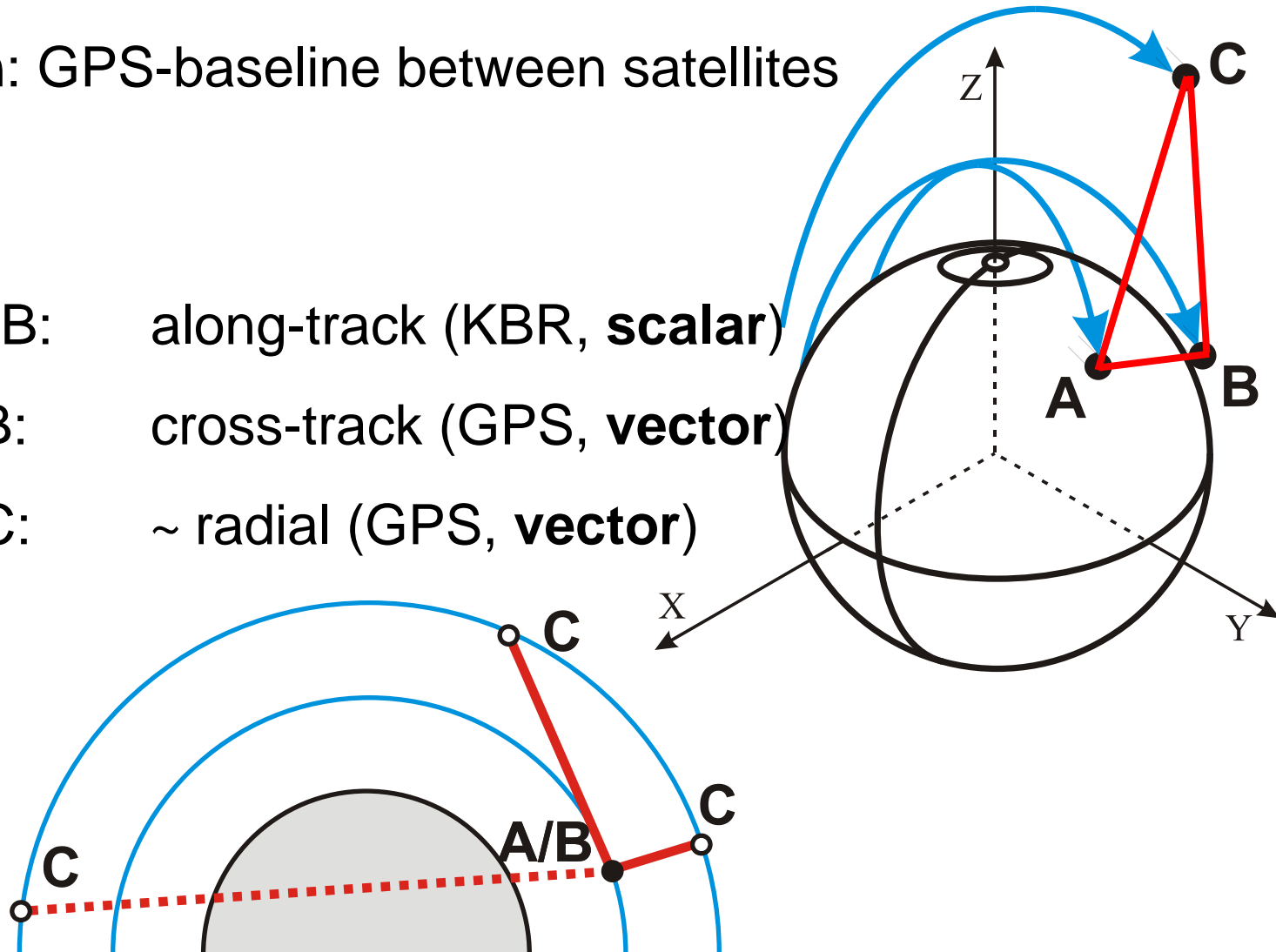
Observation: GPS-baseline between satellites

Geometry:

GRACE A-B: along-track (KBR, **scalar**)

Swarm A-B: cross-track (GPS, **vector**)

Swarm A-C: ~ radial (GPS, **vector**)



Swarm and Gravity

Use of GPS-Baseline Observations

(1) **Error propagation** using transfer coefficients for potential gradients V_x, V_y, V_z

For short baselines:

$$\Delta V / \|\Delta x\| \approx V_x$$

(2) **Full simulation** using energy integral for potential differences

$\Delta V_{\text{along-track}}$

$\Delta V_{\text{cross-track}}$

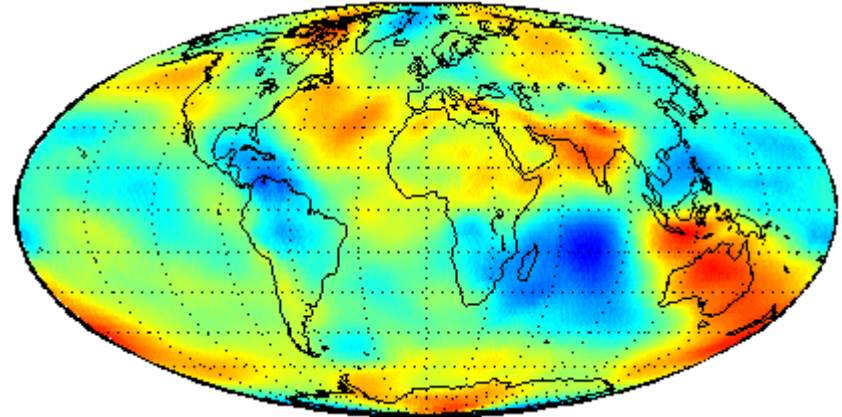
ΔV_{radial}

Swarm and Gravity

Along-Track Gradient vs. Potential Differences

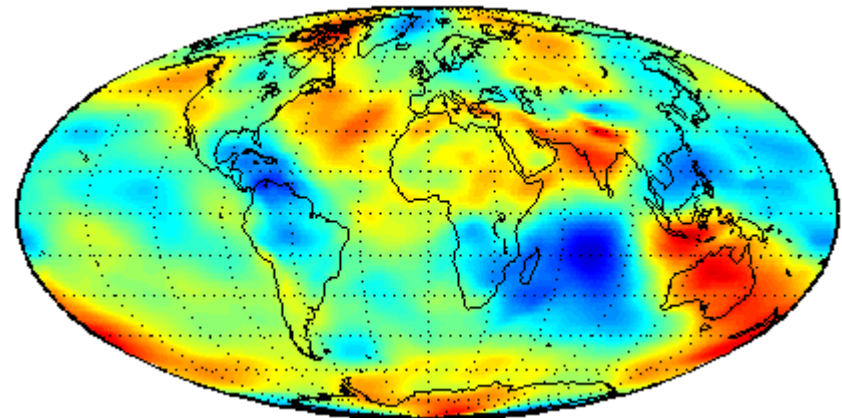
Approximate gradient

potential difference along
simulated orbit



Along-track gradient

sh-synthesis using
transfer coefficients $H_{lmk}^{\#}$



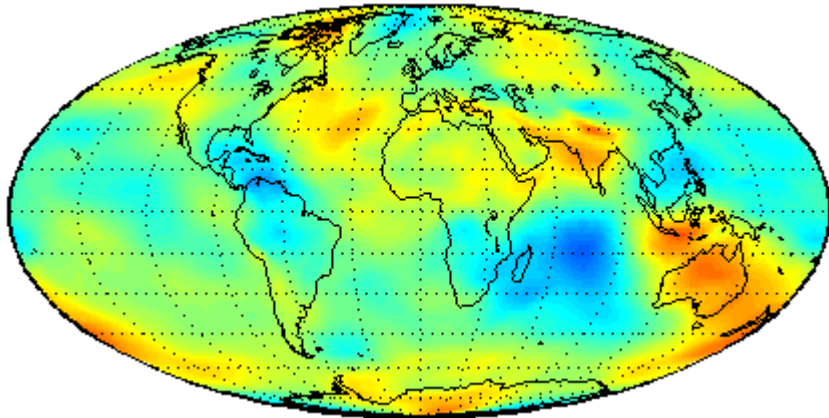
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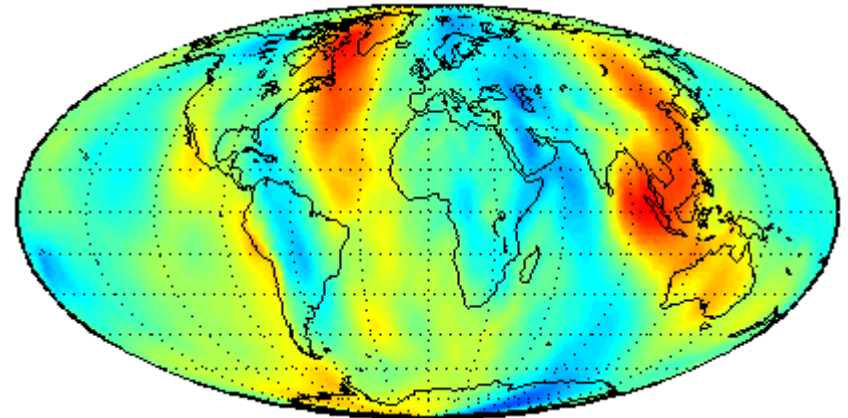
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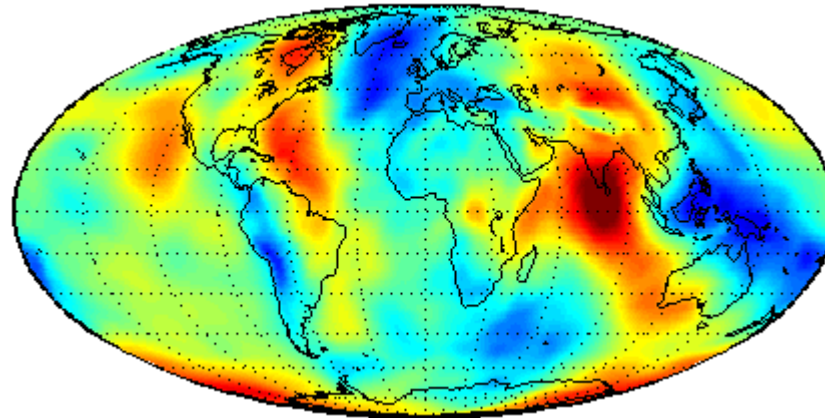
Gravity Field Gradients



along-track



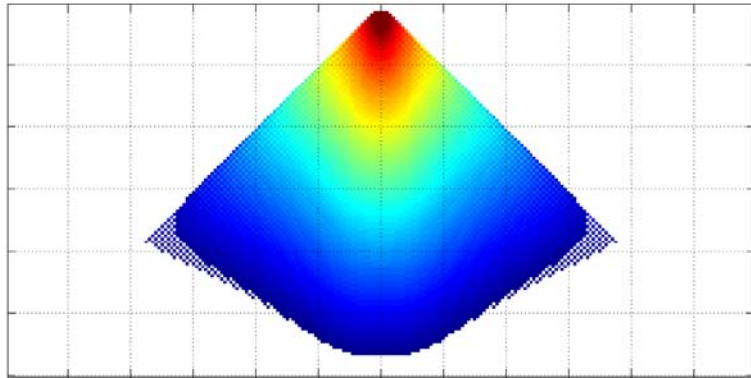
cross-track



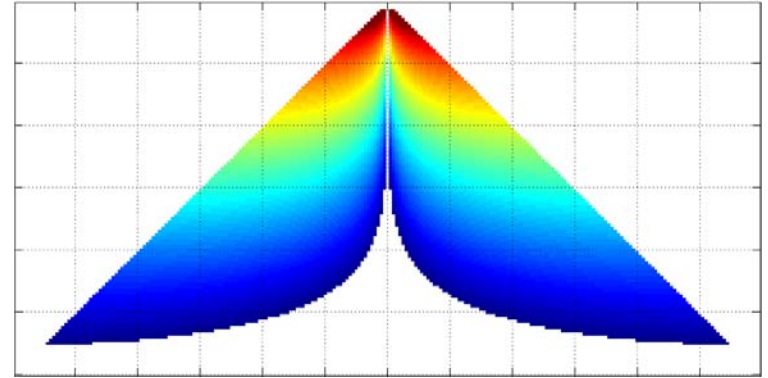
radial

SH-Error Characteristics from Different Observation Directions

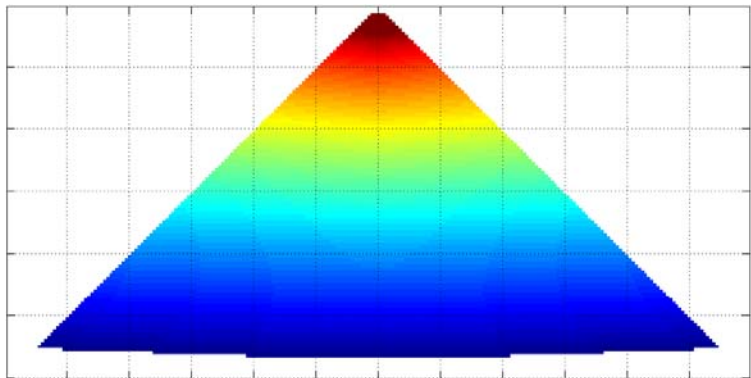
(1)



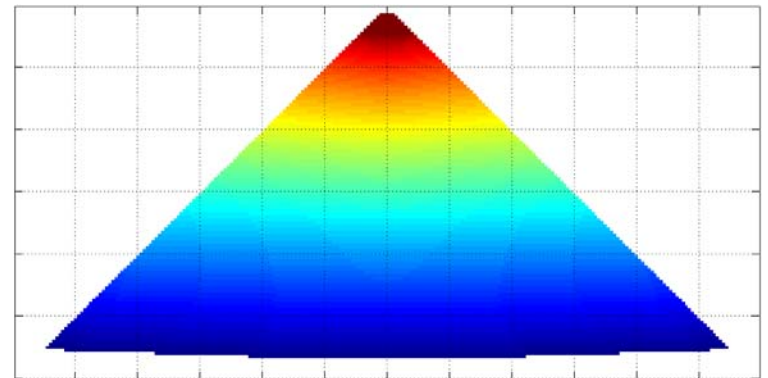
along-track



cross-track



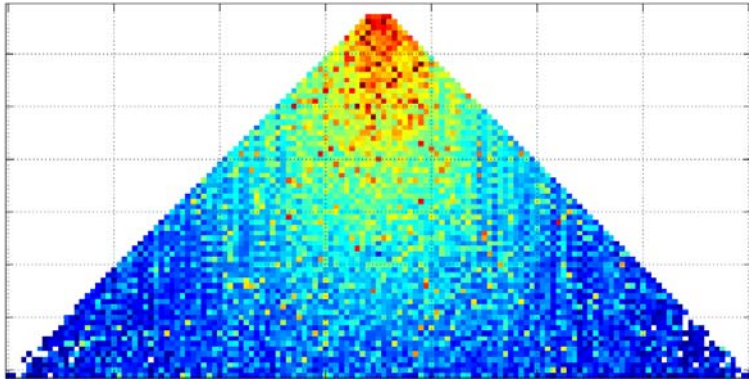
along + cross



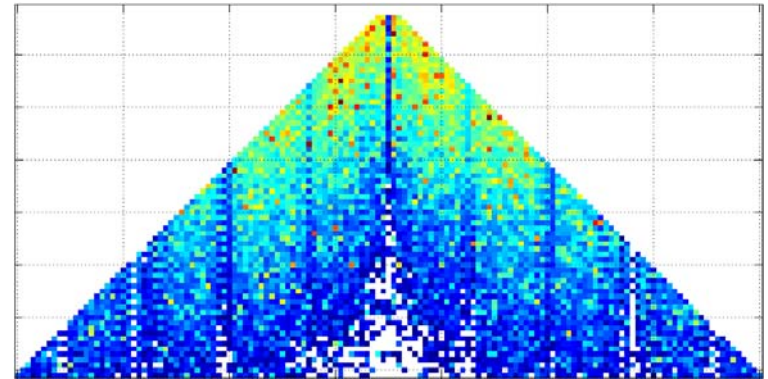
radial

SH-Error Characteristics from Different Observation Directions

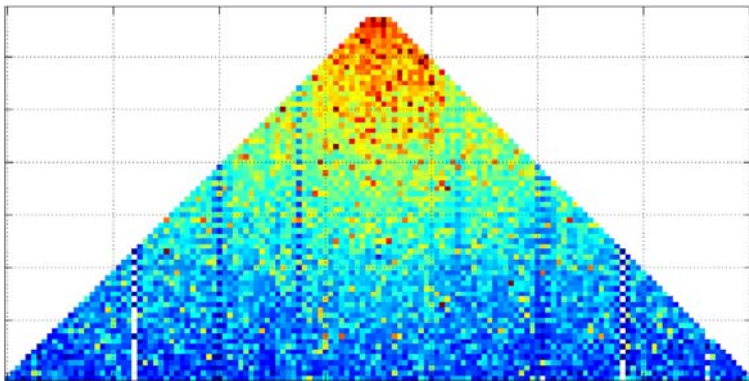
(2)



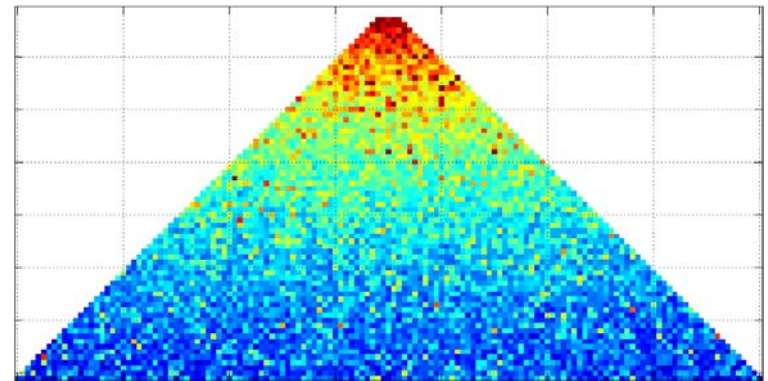
along-track



cross-track



along + cross

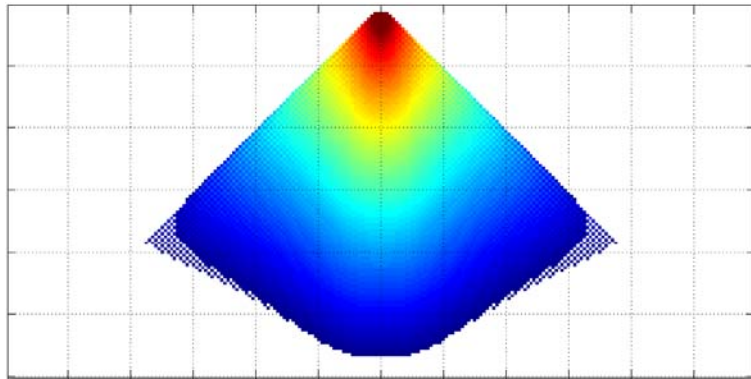


radial

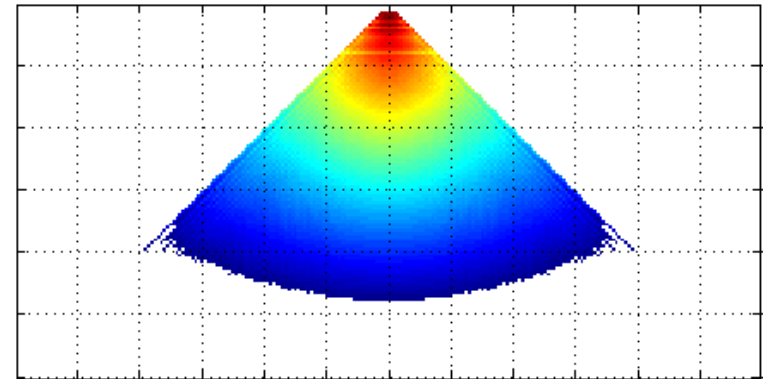
SH-Error Characteristics

Along-track simulations vs. real data (GRACE)

(1)

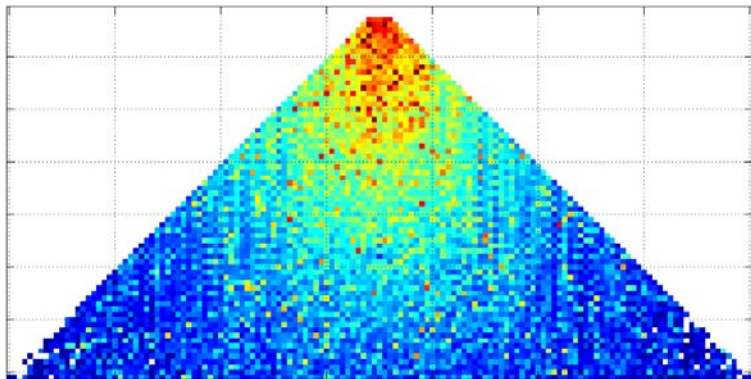


error propagation



EIGEN-GRACE02s (GFZ)

(2)



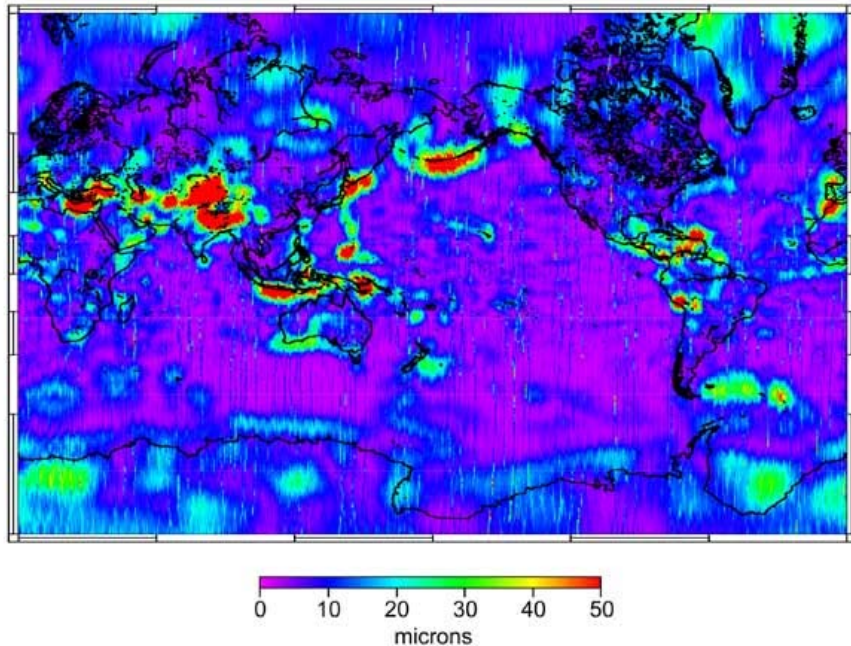
full simulation

GPS inter-satellite baseline

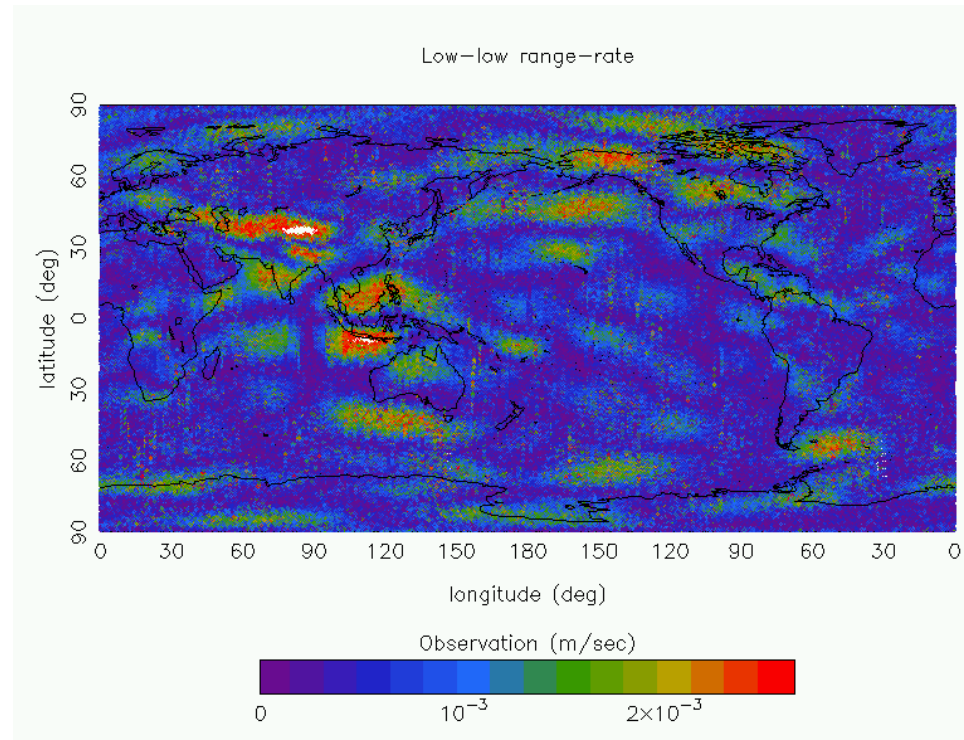
Test case: GRACE

High-pass filtered GRACE inter-satellite ranging

From KBR

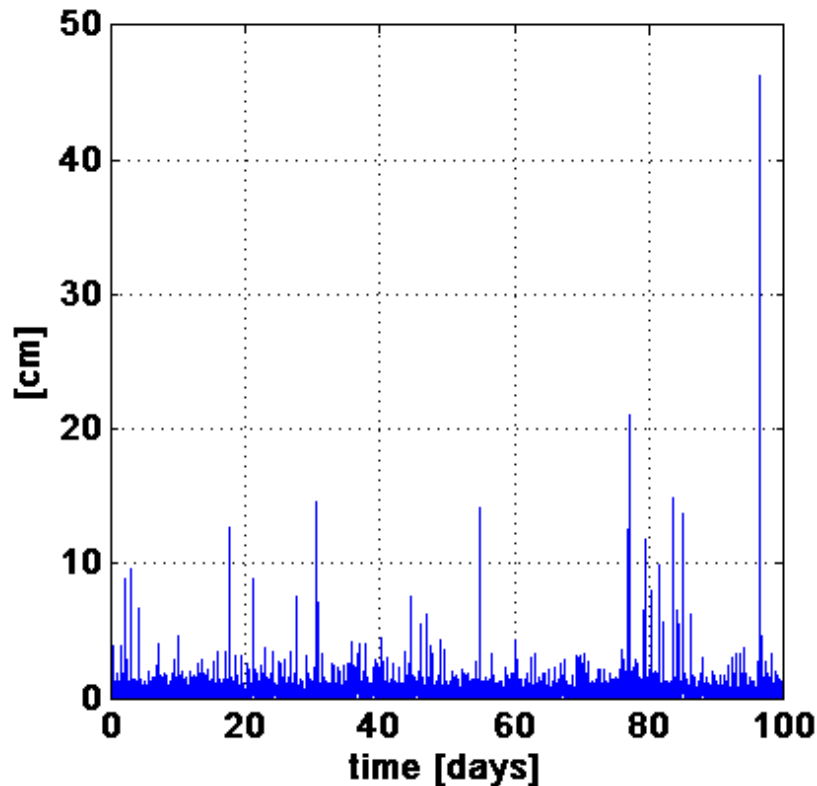


From GPS



GPS inter-satellite baseline

Test case: GRACE



Empirical range errors
(GPS vs. KBR)

$$\sigma_{\text{along}} = 2.1 \text{ mm}$$

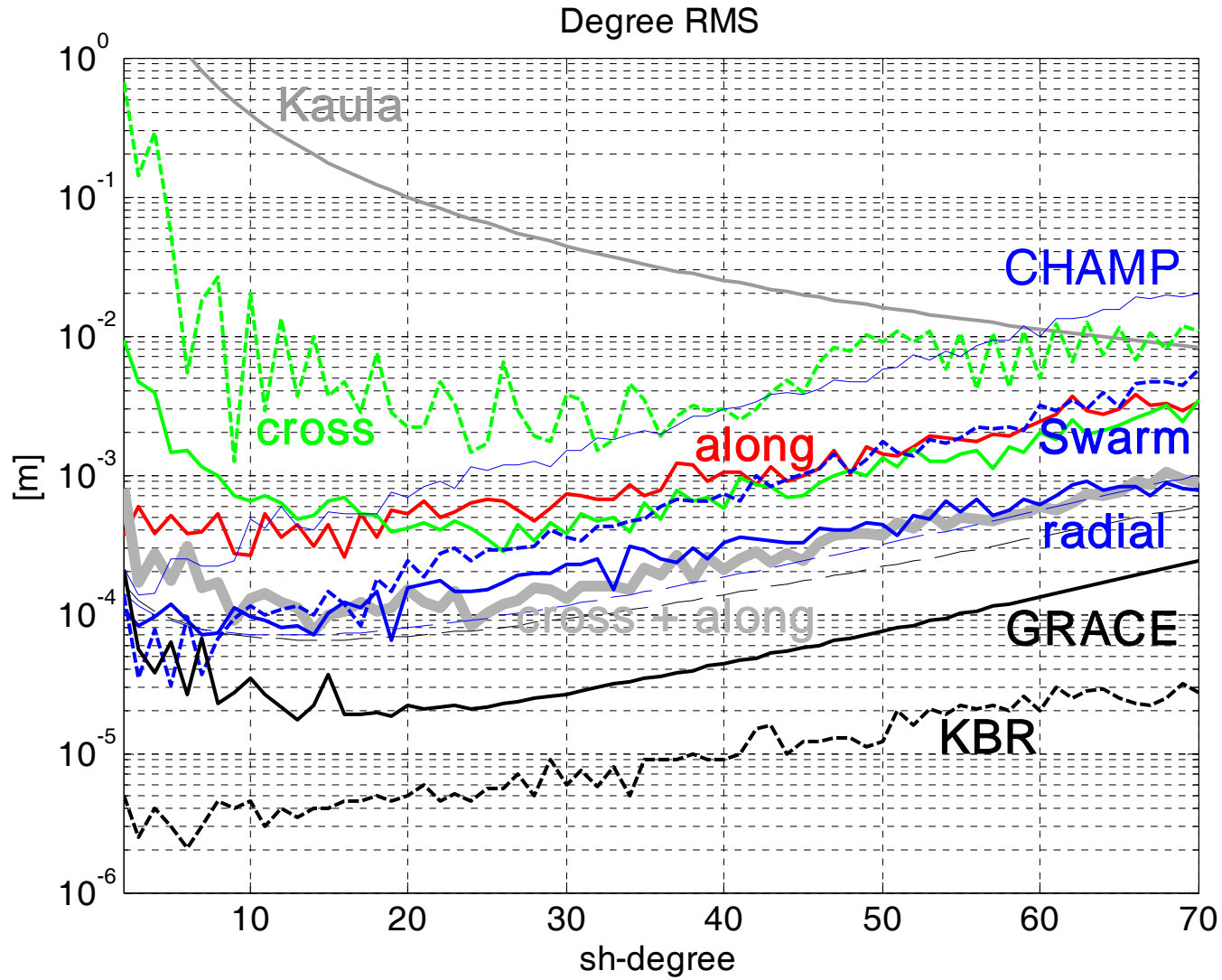
$$\sigma_{\text{cross}} = 2.0 \text{ mm}$$

$$\sigma_{\text{radial}} = 2.7 \text{ mm}$$

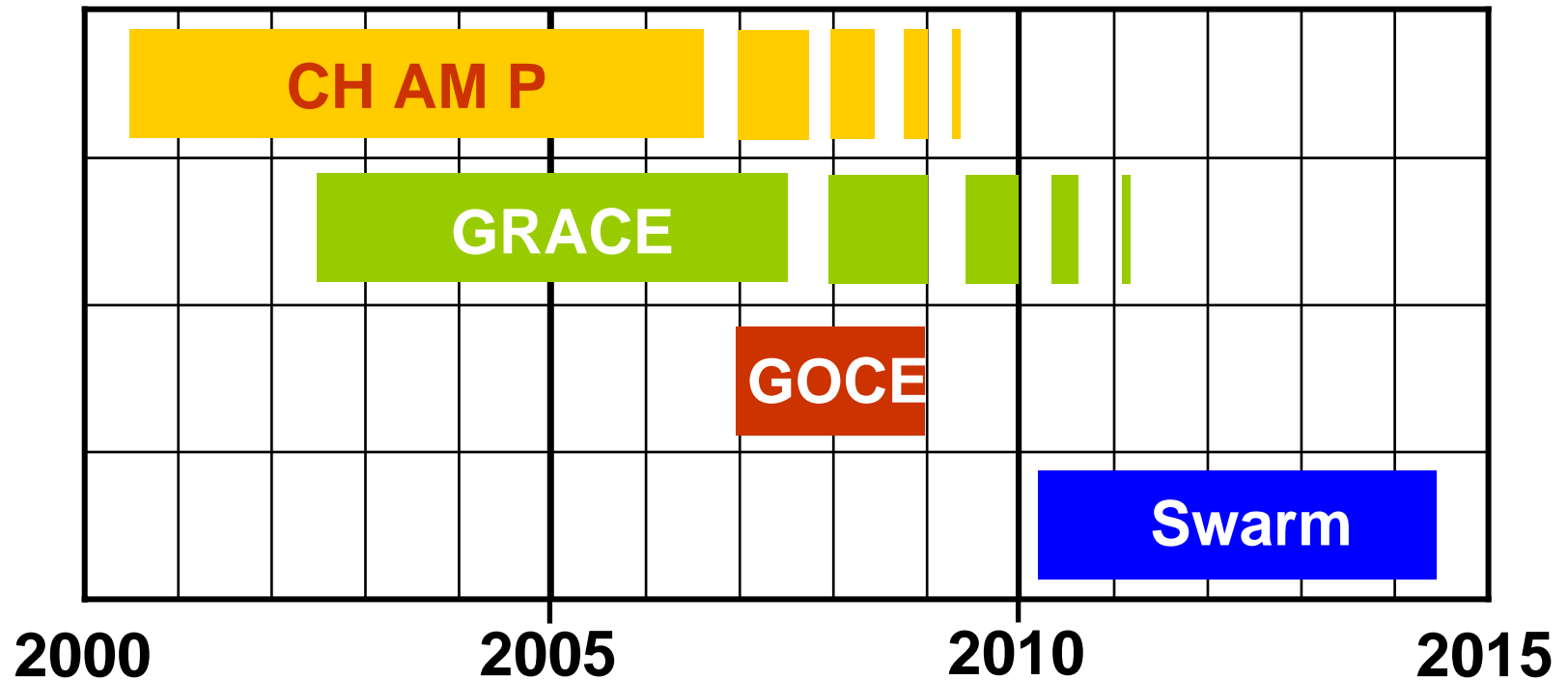
⇒ Expected range-rate accuracy:

$$\sigma \approx 10^{-2} \text{ mm/s}$$

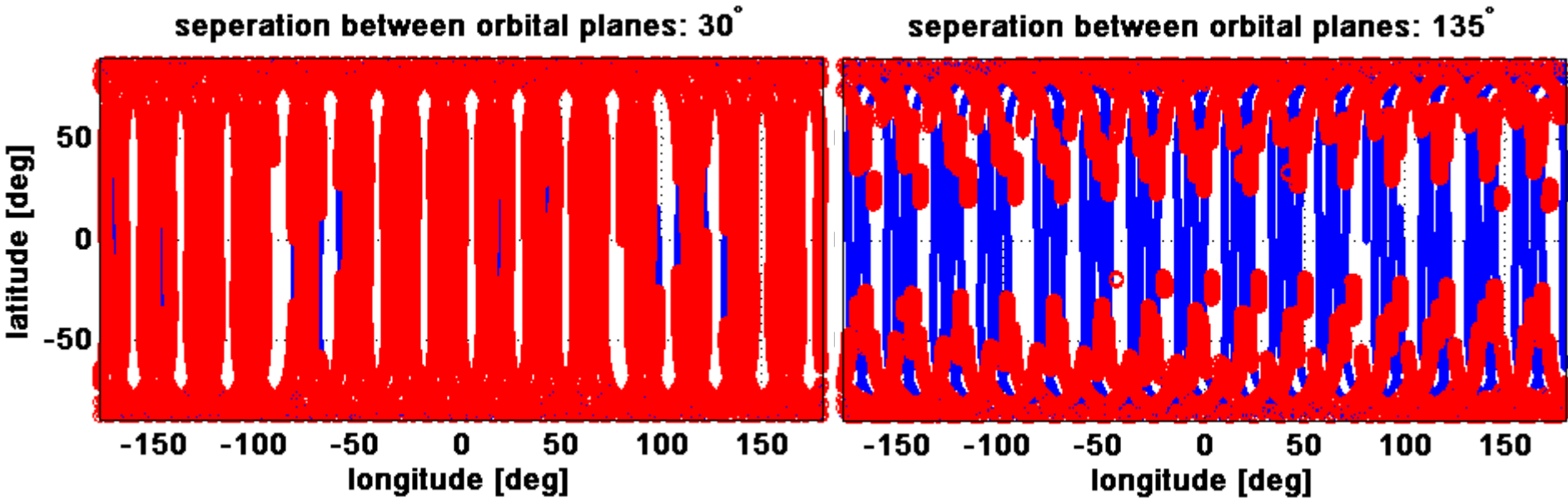
Simulation results



Time line of geopotential missions



Data distribution during the mission life time



Concluding Remarks

- **Gravity field recovery with Swarm is possible**
- **Continuation of decade of gravity field mapping possible (but, with reduced accuracy)**
- **Radial component most sensitive at end of mission (but, bad visibility)**
- **Simulation must include full GPS-simulation (visibility)**
- **Realistic error estimates for GPS?**
- **Different analysis methods possible (not necessarily energy integral)**