

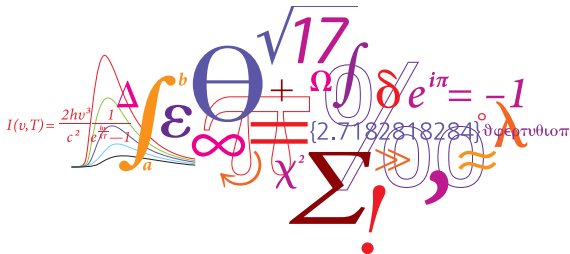
Gyre-driven decay of the Earth's magnetic dipole

Chris Finlay¹, Julien Aubert² & Nicolas Gillet³

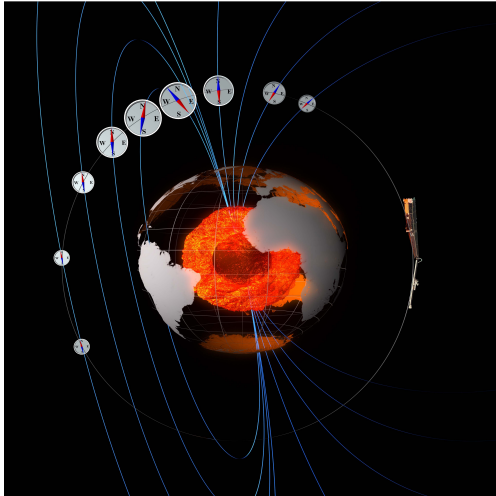
1: DTU Space, Technical University of Denmark

2: IPGP, Université Sorbonne Paris Cité

3: ISTerre, l'Université Joseph Fourier, Grenoble

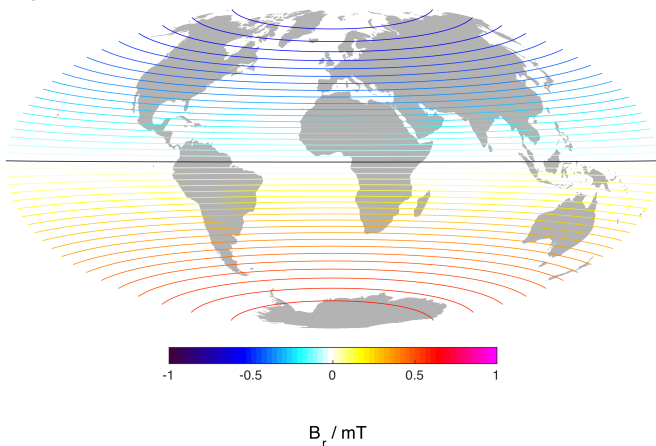


Earth's magnetic field: Predominantly an axial dipole



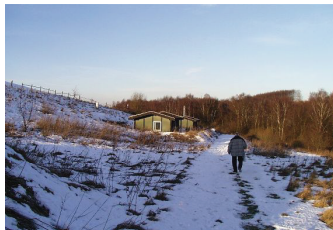
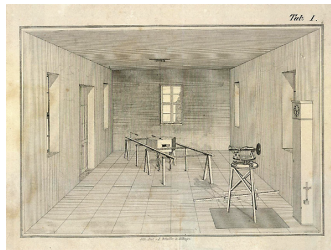
Credit:ESA

Axial dipole field

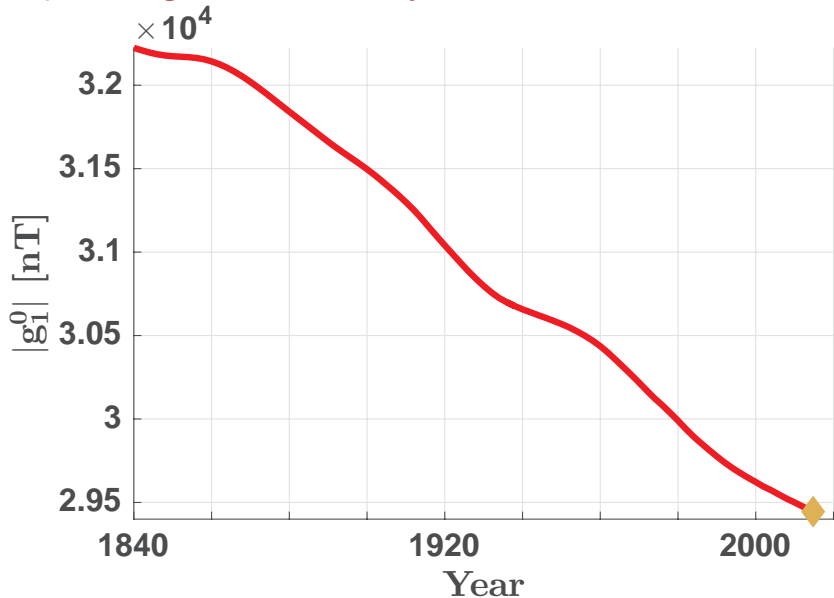


- ▶ Axial dipole g_1^0 is the first term in the spherical harmonic expansion of \mathbf{B} .
- ▶ Dipole represents 99.9% of field energy at $10 R_e$.
- ▶ Dipole represents 93% of field energy at Earth's surface, $1 R_e$.
- ▶ Dipole represents 37% of field energy at the core surface, $0.55 R_e$

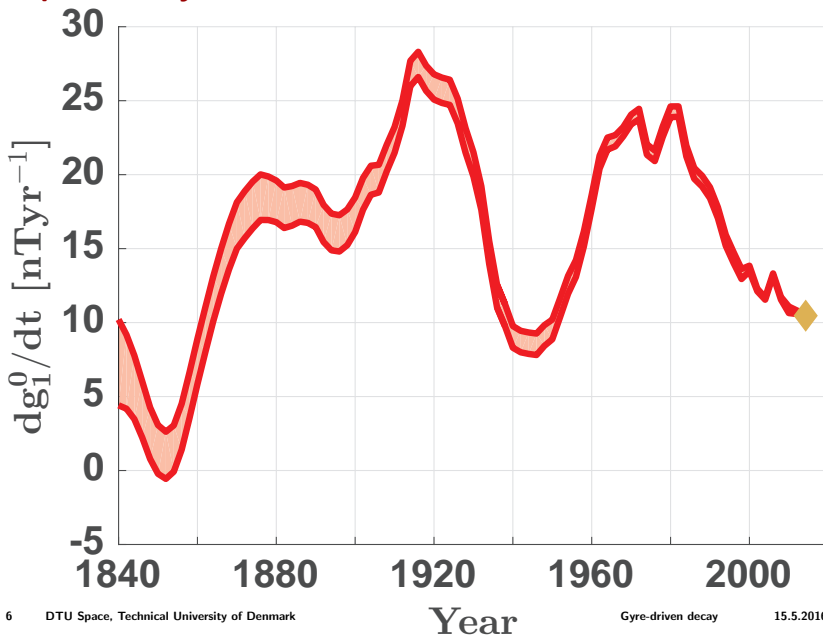
175 years of absolute observations



Dipole magnitude has decayed $\sim 10\%$ since 1840



Dipole decay rate fluctuations on decadal time scales



Origin of the decay: MHD processes in the core

The rate of change of the dipole moment is

$$\frac{d\mathbf{m}}{dt} = \frac{1}{2} \int \hat{\mathbf{r}} \times \frac{\partial \mathbf{J}}{\partial t} dV = \frac{3}{2\mu_0} \int \frac{\partial \mathbf{B}}{\partial t} dV. \quad (1)$$

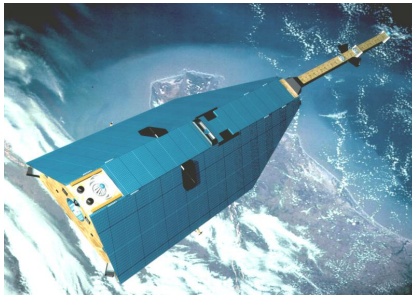
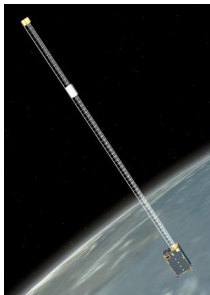
Substituting from the magnetic induction equation

$$\frac{d\mathbf{m}}{dt} = \frac{3}{2\mu_0} \int [\nabla \times (\mathbf{u} \times \mathbf{B}) + \eta \nabla^2 \mathbf{B}] dV. \quad (2)$$

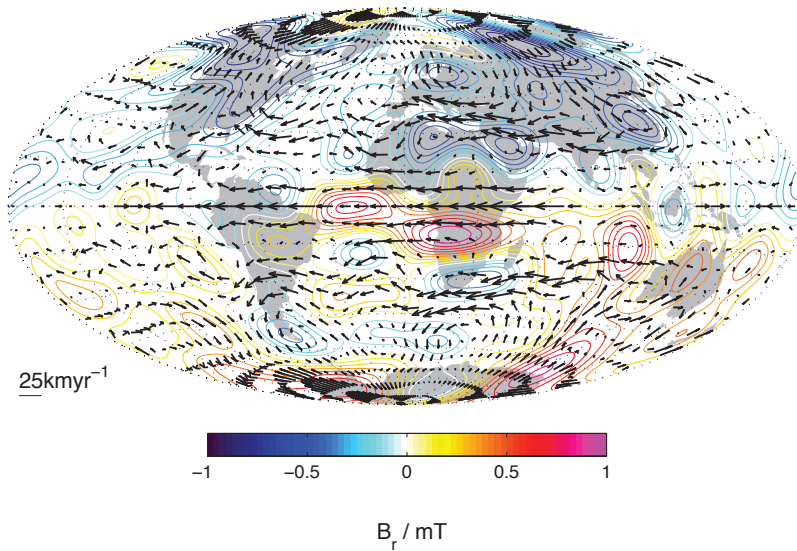
Taking the axial component, expanding and integrating by parts, find that the axial dipole moment (ADM) change can be written as

$$\frac{dm_z}{dt} = \underbrace{-\frac{3}{2\mu_0} \int u_\theta \sin \theta B_r dS}_{\text{ADM change due to meridional advection}} + \underbrace{\frac{3\eta}{2\mu_0} \int \hat{\mathbf{z}} \cdot \nabla^2 \mathbf{B} dV}_{\text{ADM change due to Ohmic diffusion}}. \quad (3)$$

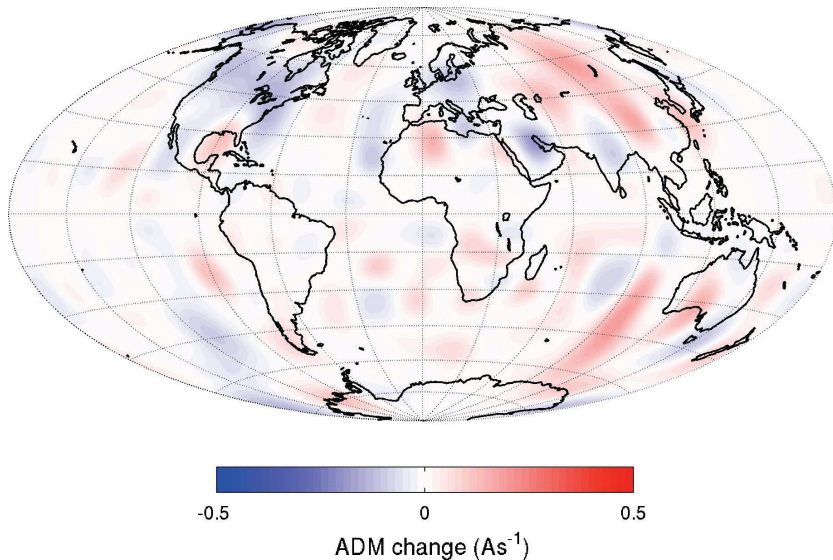
Satellite vector field observations 1999-2010



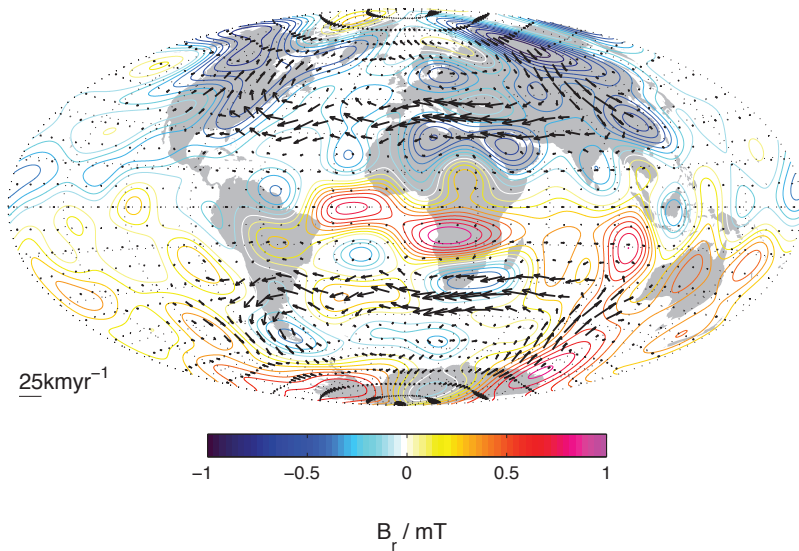
Core surface field and flow: 2000-2010



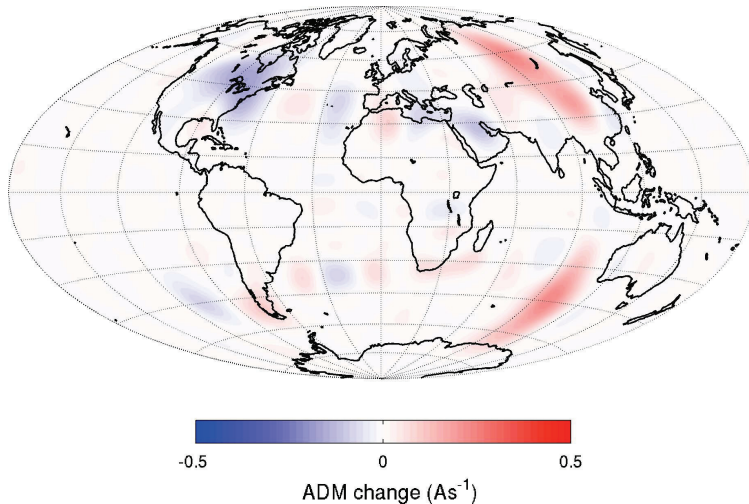
Contributions from meridional flux transport decay



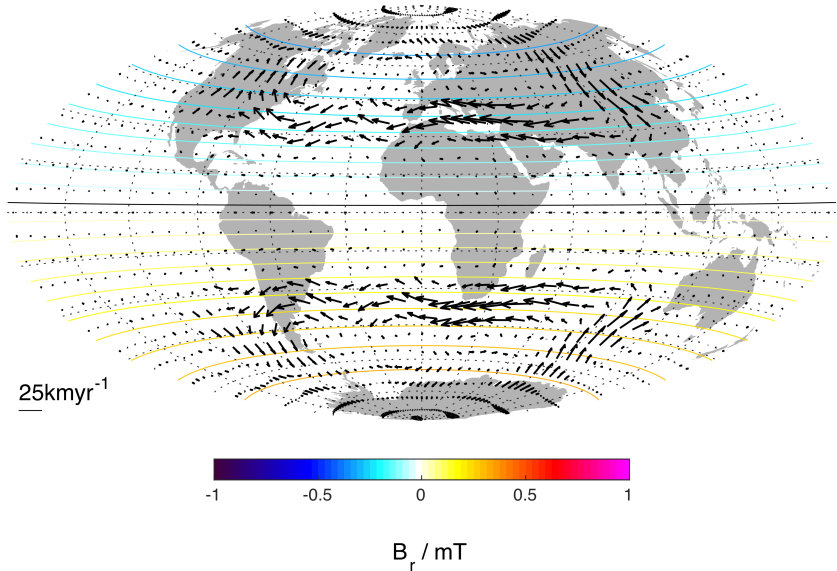
Simplified model: Observed field plus filtered gyre



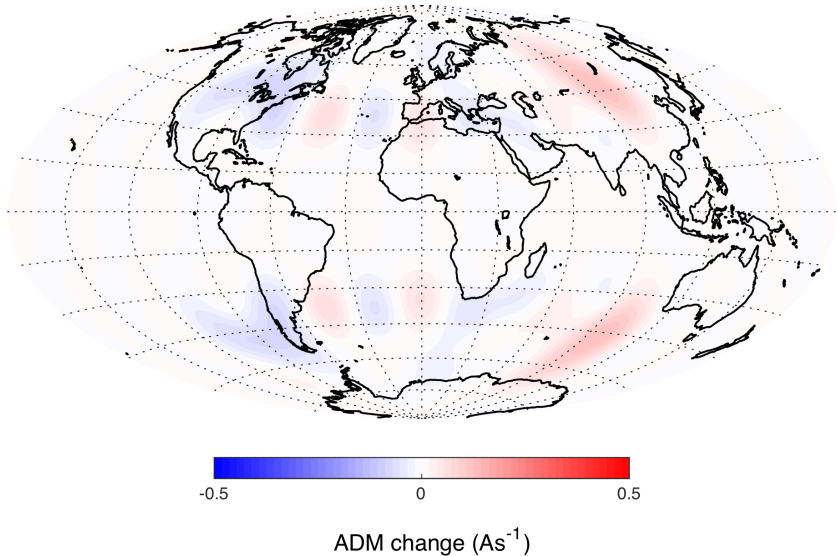
Simplified model: Observed field plus filtered gyre



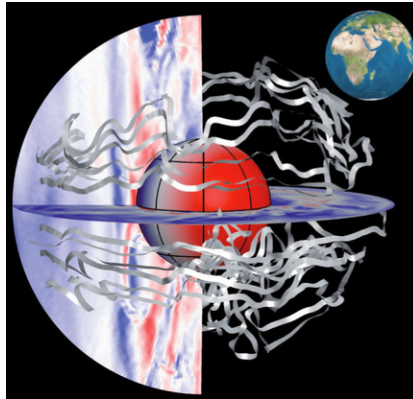
Comparison with case of pure dipole field



Comparison with case of pure dipole field

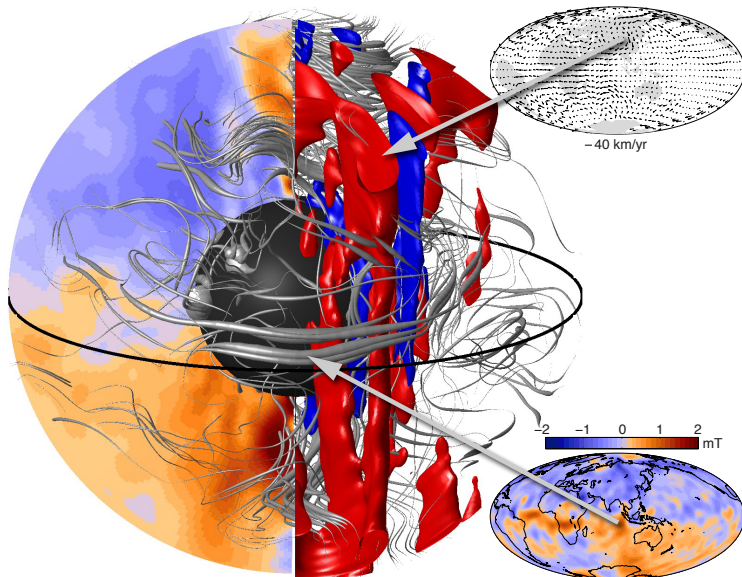


A 3D dynamo model including magnetic diffusion

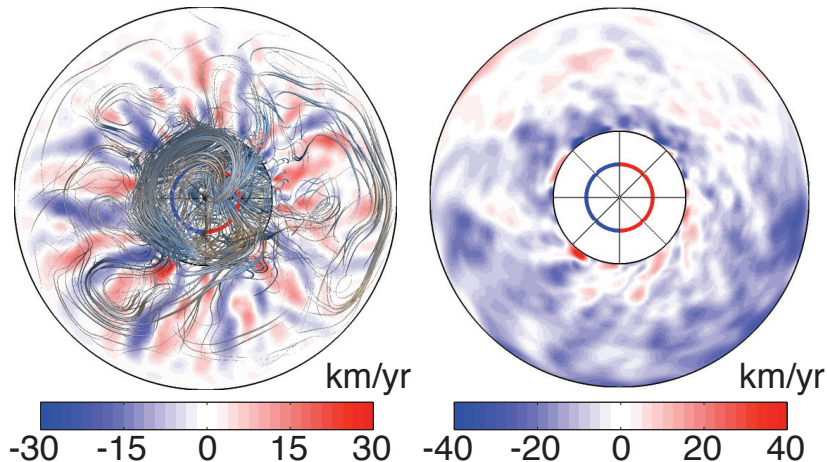


- ▶ Coupled Earth Dynamo Model ([Aubert et al., 2013](#))
- ▶ EM coupling at ICB, gravitational coupling btw IC and mantle
- ▶ Relatively high magnetic Reynolds number, $Rm = 942$
- ▶ Produces a planetary-scale gyre as in core flow inversions

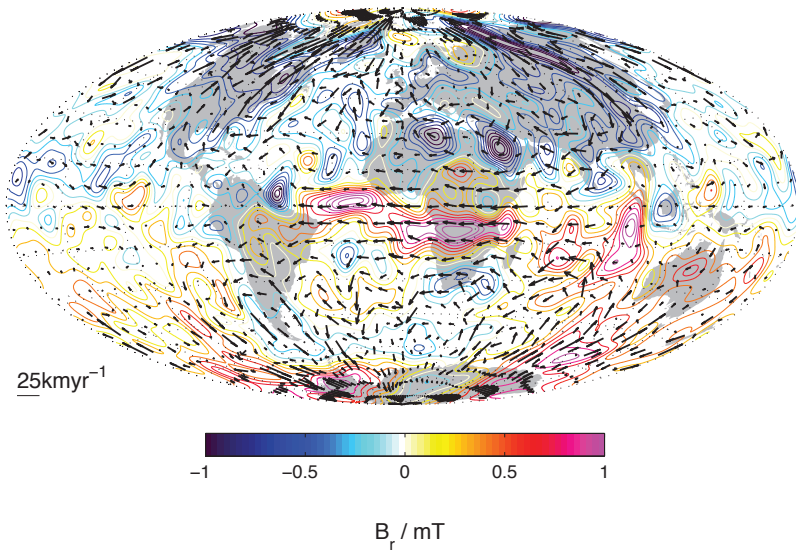
Estimated geodynamo state in 2015



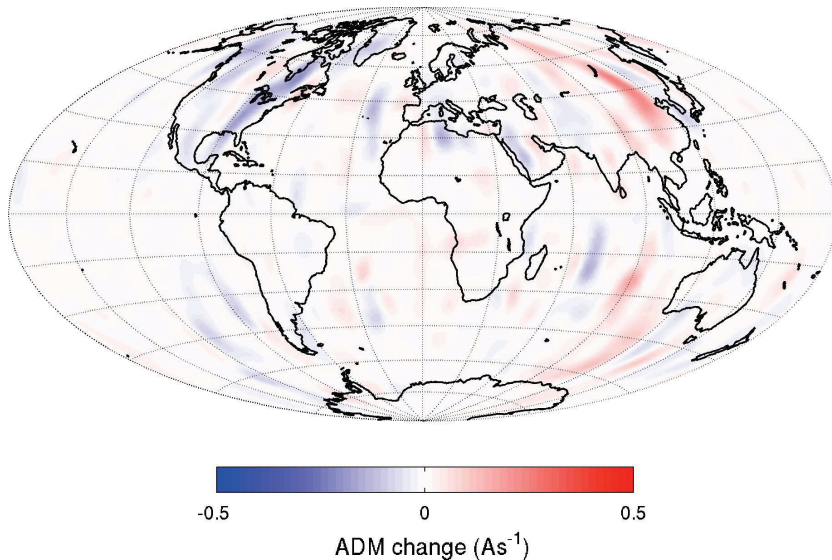
Estimated field and flow within core in 2015



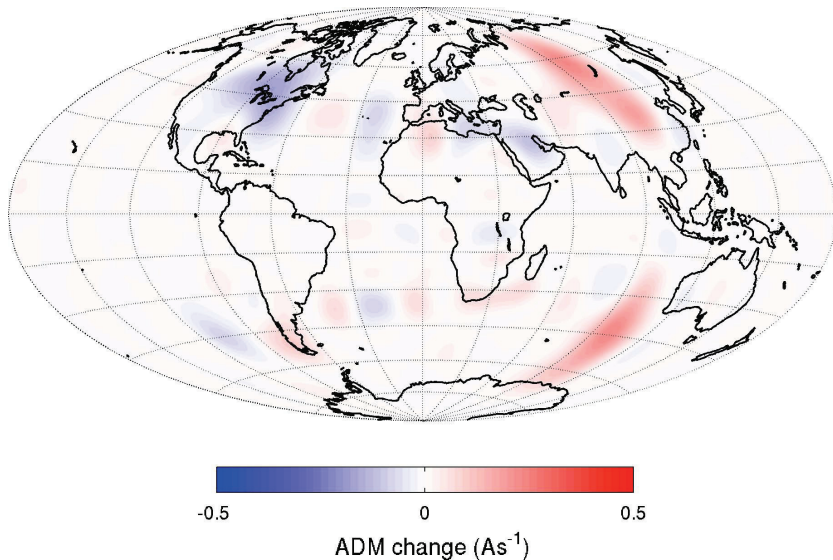
Estimated core surface field and flow in 2015



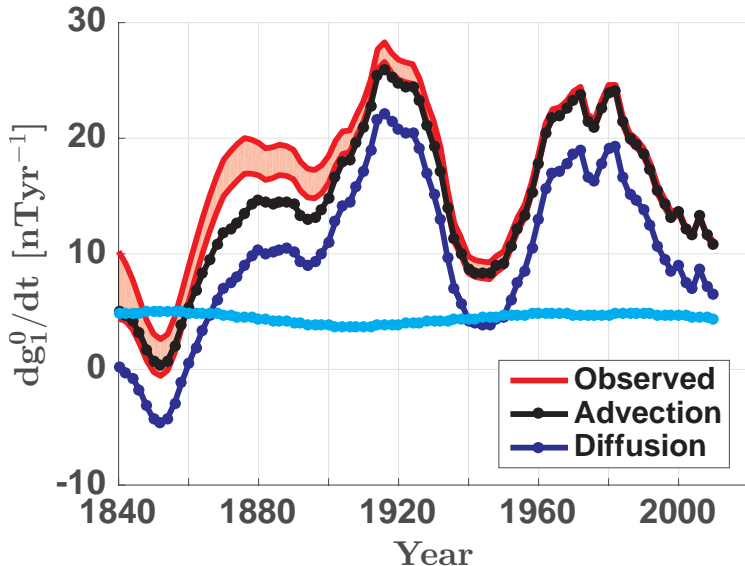
Dipole decay due to meridional flux transport in 2015



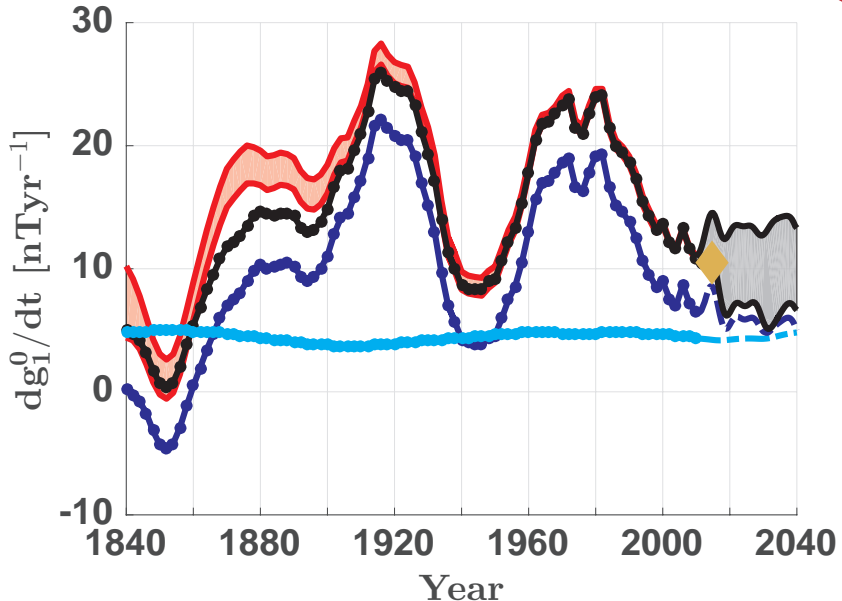
Comparison with simple gyre model



Retrospective analysis: Contributions to historical dipole decay

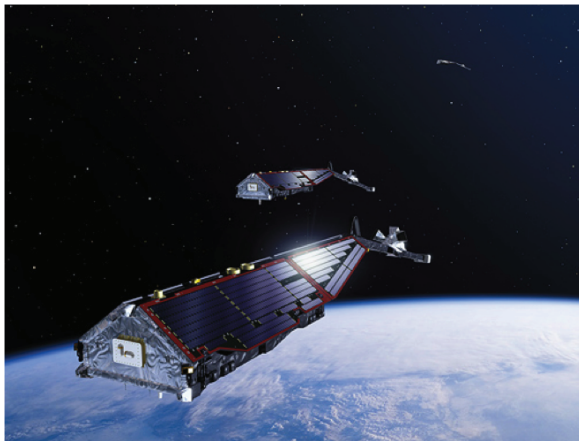


Forecast of dipole decay rate for upcoming decades



- ▶ The Earth's magnetic dipole has been decaying for the past 175 yrs.
- ▶ Planetary gyre transporting normal flux equatorward and reversed flux poleward accounts for much of the decrease.
- ▶ Mechanism requires field asymmetric (e.g. South Atlantic Anomaly).
- ▶ Dipole decay fluctuations due to fluctuations in meridional flux transport by the gyre.
- ▶ Magnetic diffusion makes a secondary, almost steady contribution.
- ▶ Mechanism suggests dipole will continue to decay, at least for next few decades.

Outlook: Using *Swarm* data for more detailed tests



- Do fluctuations in equatorward flow correlate with fluctuations in the decay rate?

[PhD thesis of M. Hammer]

Inverse geodynamo modelling: A Kalman filter approach

[See Aubert (2013, 2014) for full details]

- ▶ (i) Estimate \mathbf{B} throughout core \mathbf{x}_B to degree 30, from the poloidal field at the surface to degree 13, using the a-priori covariance matrix \mathbf{P}_B derived from a large number of states \mathbf{x}_B :

$$\mathbf{x}_B = \mathbf{P}_B \mathbf{H}_B^T (\mathbf{H}_B \mathbf{P}_B \mathbf{H}_B^T + \mathbf{R}_B)^{-1} \mathbf{b} \quad \text{where} \quad \mathbf{H}_B \mathbf{x}_B = \mathbf{b} \quad (4)$$

- ▶ (ii) Estimate the core surface flow \mathbf{x}_{fs} using the poloidal SV from a field model corrected by diffusion estimates from (i), $\dot{\mathbf{b}}$:

$$\mathbf{x}_{fs} = \mathbf{P}_{fs} \mathbf{M}^T (\mathbf{M} \mathbf{P}_{fs} \mathbf{M}^T + \mathbf{R}_{\dot{B}})^{-1} (\dot{\mathbf{b}} + \mathbf{c}) \quad \text{where} \quad \mathbf{M} \mathbf{x}_{fs} = \dot{\mathbf{b}} + \mathbf{c} \quad (5)$$

- ▶ (iii) Estimate velocity throughout core \mathbf{x}_u from \mathbf{x}_{fs} using the a-priori covariance matrix \mathbf{P}_u

$$\mathbf{x}_u = \mathbf{P}_u \mathbf{H}_u^T (\mathbf{H}_u \mathbf{P}_u \mathbf{H}_u^T)^{-1} \mathbf{x}_{fs} \quad \text{where} \quad \mathbf{H}_u \mathbf{x}_u = \mathbf{x}_{fs} \quad (6)$$

- ▶ Provides an estimate of (\mathbf{u}, \mathbf{B}) throughout the core, at a given epoch, from a field model plus prior statistics from a self-consistent 3-D numerical dynamo simulation.

