The X-ray monitor JEM-X

Purpose:

Precise position determination (< 1 arcmin)

Cover low energyspectrum 3 - 30 keV

Properties:

Field-of-view 12 deg diameter

5 deg FCFOV

Energy resolution 16% @ 7 keV

10% @ 20 keV

Imaging: Niels J. Westergaard, Niels Lund, Carl Budtz-Jørgensen

Danish Space Research Institute

Spectra: Stefan Larsson

Stockholm University

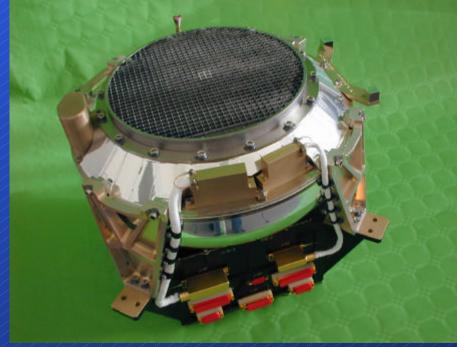
INTEGRAL with JEM-X





JEM-X HARDWARE





Modes of operation

Full imaging FULL

Restricted imaging REST

Rudimentary energy binning, limited time res.

Non-imaging formats (all detector):

Spectral format SPEC

Timing format TIME

Spectral timing SPTI

Grey filter event rejection

When the event buffer is filled above certain level a random rejection of events will be initiated autonomously

Level G will reject a fraction (31 - G)/32

When buffer filling is low again the grey filter level will increase again.

Image production

```
Steps in the process:
```

Level Function

BIN_I Shadowgram building
BKG Background subtraction
IMA Image reconstruction

IMA Flatfielding

IMA Source finding

IMA Vignetting correction

IMA2 Mosaicking

Executable

j_ima_shadowgram

(included)

j_ima_basic_recon

(included)

j_ima_src_find

j_ima_cor_intensity

j_ima_mosaic (OSA5)

Shadowgram building

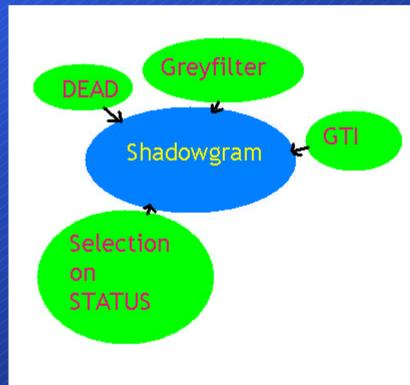
Event selection based on STATUS flag (rowSelect)
Correction for deadtime
Correction for grey filter

```
TELAPSE = Tstart - Tstop

ONTIME = GTI<sub>1</sub> + GTI<sub>2</sub> + ...

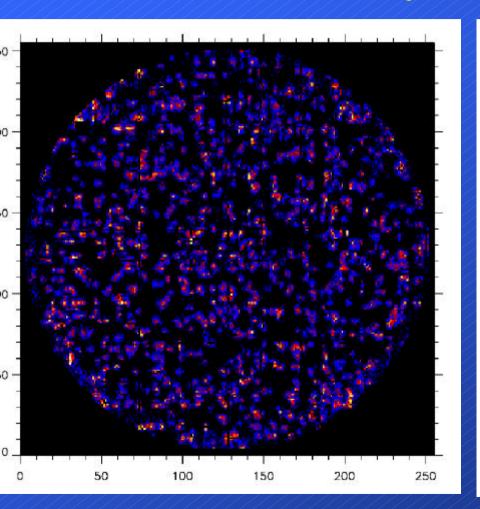
EXPOSURE = ONTIME * corfactor
```

shdType = 0 'skew' shadowgram shdType = 1 standard shadowgram

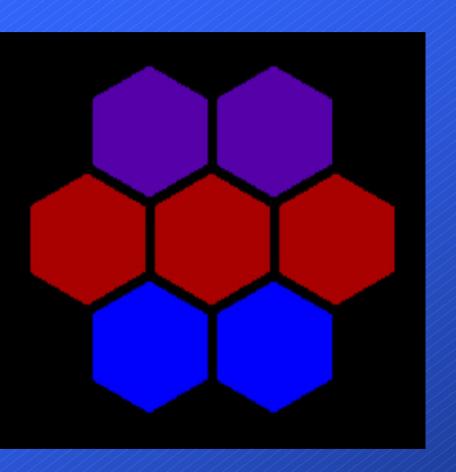


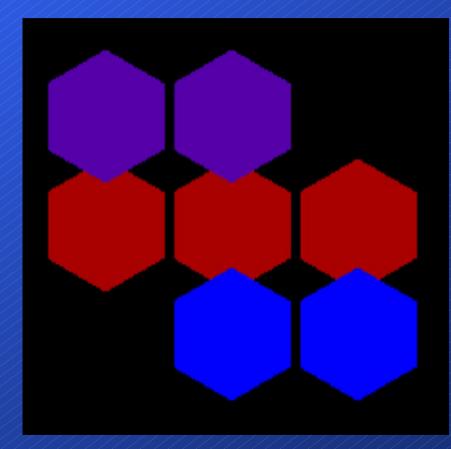
Mask pattern

25% HURA - but almost not cyclic, hence ghosts almost eliminated



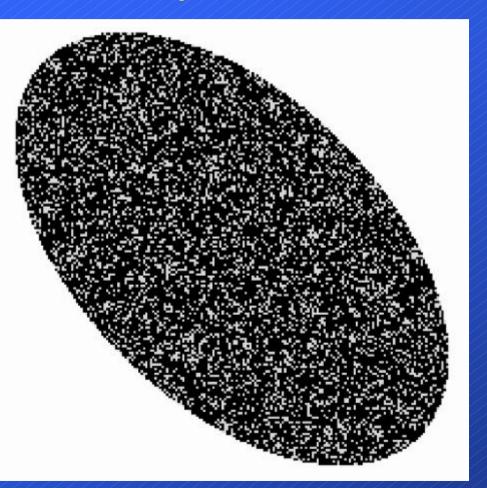
The soon to disappear skew shadowgram



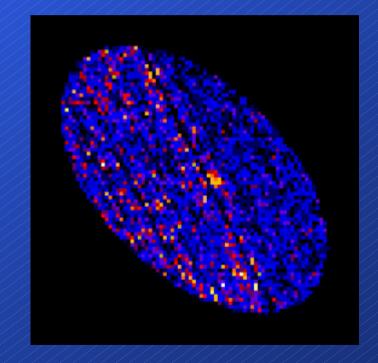


Mask and shadowgram

Mask array from IMOD GRP



Shadowgram from OSA4 shdType=0



Algorithm

Balanced correlation

For each relative position of detector and mask: Count open and closed mask cells 'on the detector' (N_{open} and N_{closed}) Sum number of counts corresponding to open and closed cells

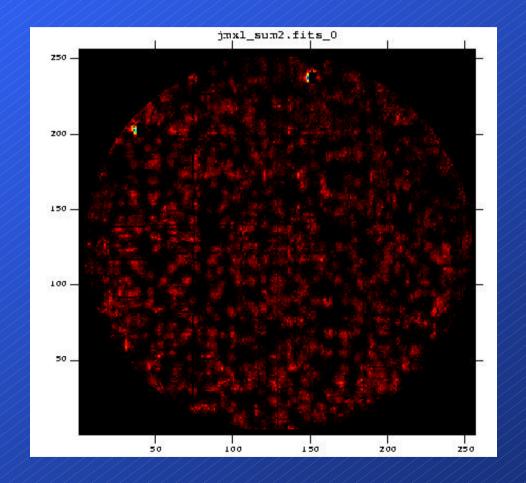
$$t = \frac{N_{\text{open}}}{N_{\text{open}} + N_{\text{closed}}} \tag{1}$$

Balanced reconstruction:

$$I = (1 - t) \sum S_{\text{open}} - t \sum S_{\text{closed}}$$
 (2)

Leaking calibration sources

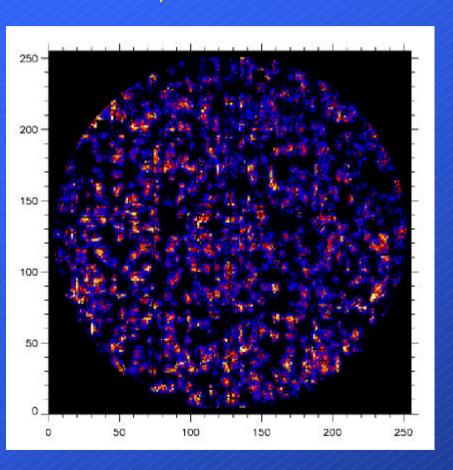
Shadowgram Crab Nebula on axis Rev 102 3 - 6 keV

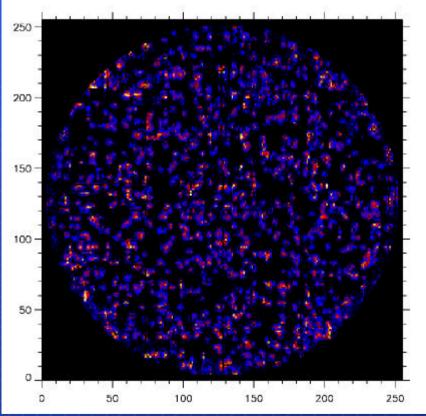


Shadowgram low, medium energy

3-6 keV, leak removed.

6-12 keV, leak removed

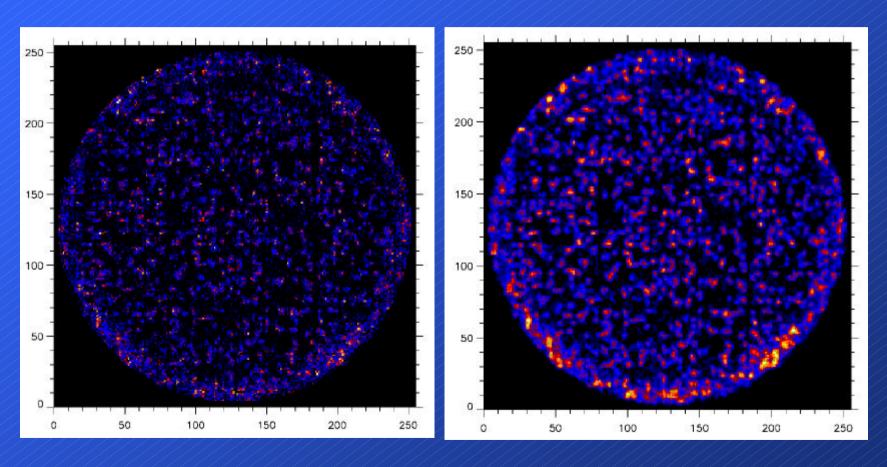




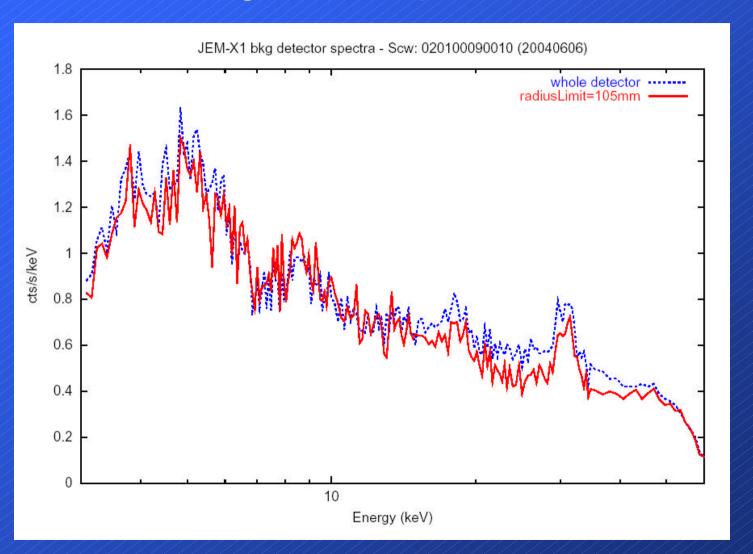
Shadowgram, high energy

12-25 keV

Smoothed version



Background spectrum

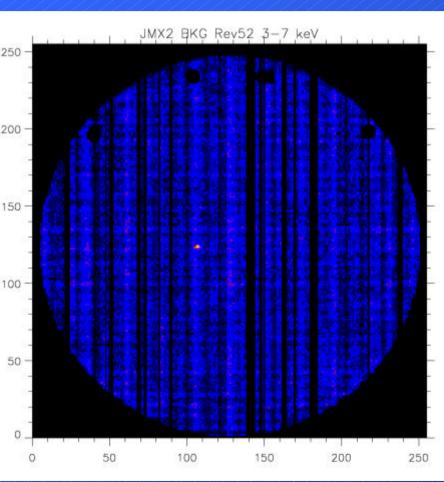


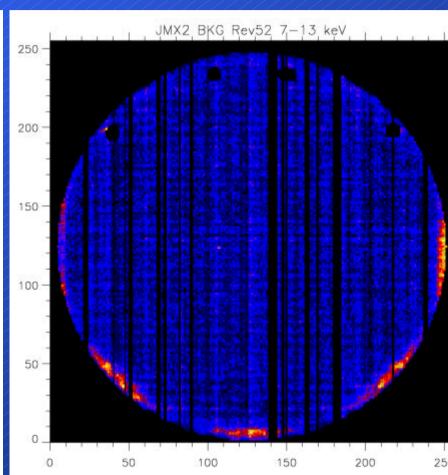
Background (1)

3 - 7 keV

30,000 s

7 - 13 keV

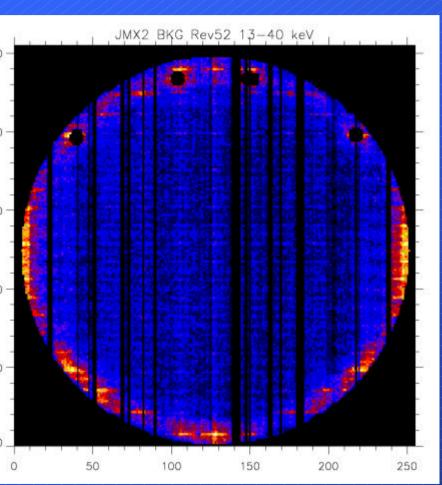


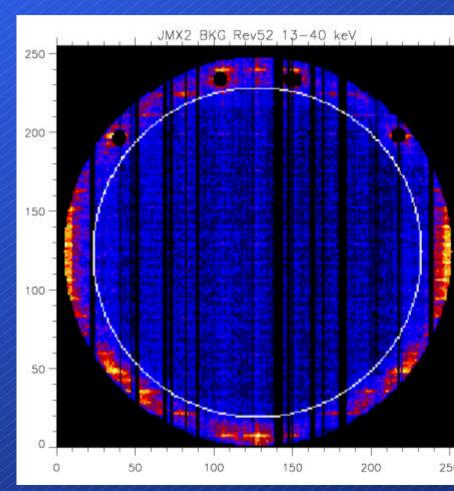


Background (2)

13 - 40 keV

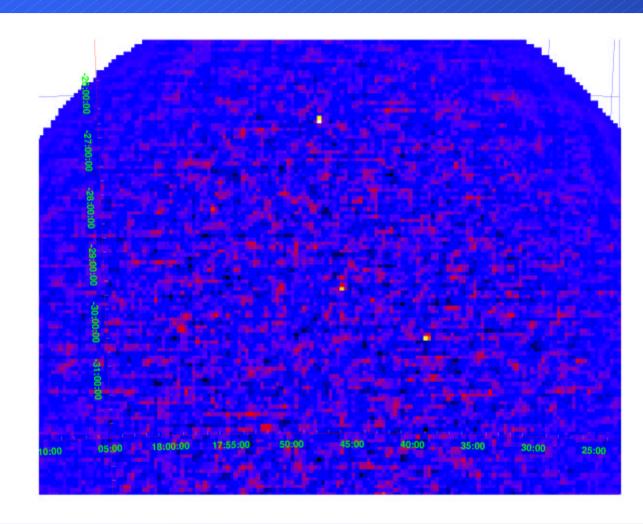
30,000 s radiusLimit = 105





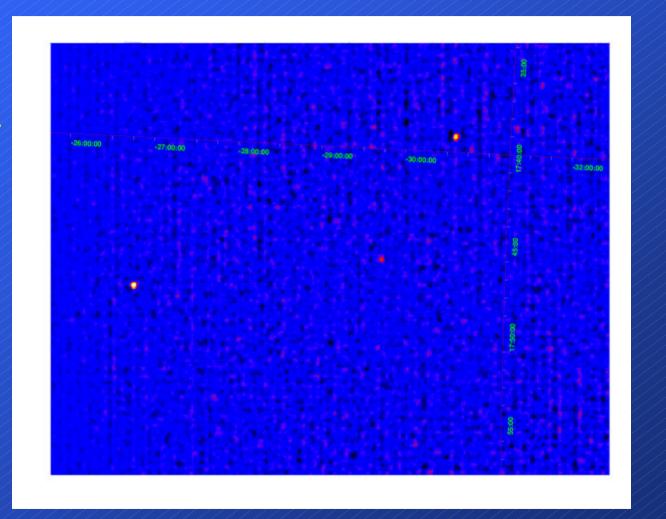
Basic reconstruction example

OSA4 result
Rev 53, PID 66
6-10 keV
Galactic Center
region



Fine resolution example

Rev 53 PID 66 Galactic center 6-10 keV



Vignetting function

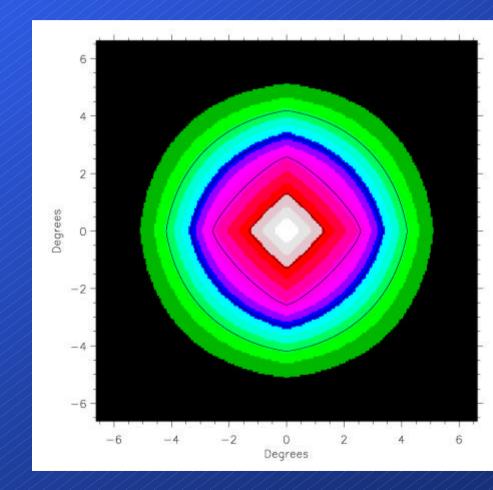
Vignetting array from IMOD-GRP
This is the average vignetting

BUT

complicated geometry between mask and collimator and detector

More work!!

Contours: 20, 40, 60, and 80%



New member of ISSW family

Midisky - to become j_ima_iros with OSA5

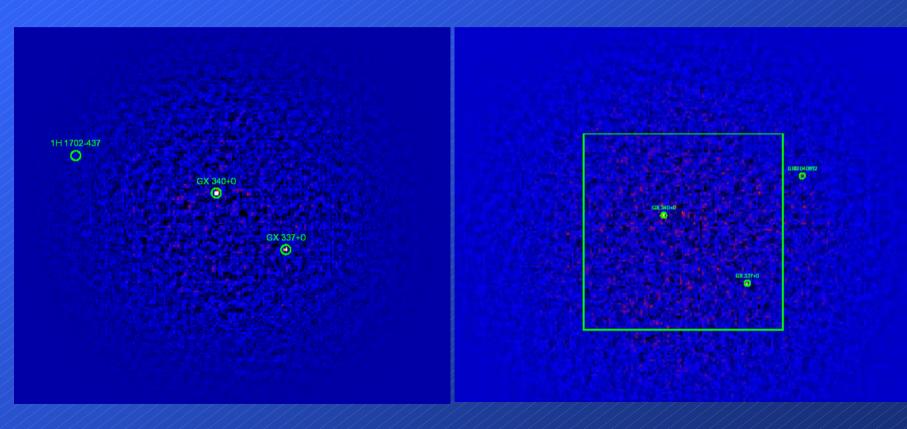
- Backprojection for source finding
- Iterative removal of sources
- As complete modelling of detector as possible including collimator imperfections
- The spectrum and lightcurve extraction based on this modelling will be included in the ISSW as well

Now: ftp.dsri.dk/pub/jemx/midisky/midi/midi_distribution_2.tar.gz

Midisky example GRB040812

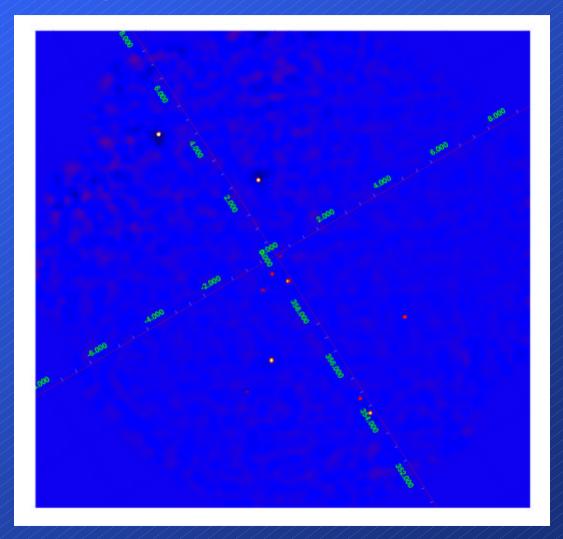
Full science window

20 s during burst



Midisky manipulated mosaic

In IROS each source gets a position and a strength. Here they are taken out, the remainder smoothed; then the sources have been put back in.



Fin

After running jemx_science_analysis up to and including level IMA you'll find in the .../res directory:

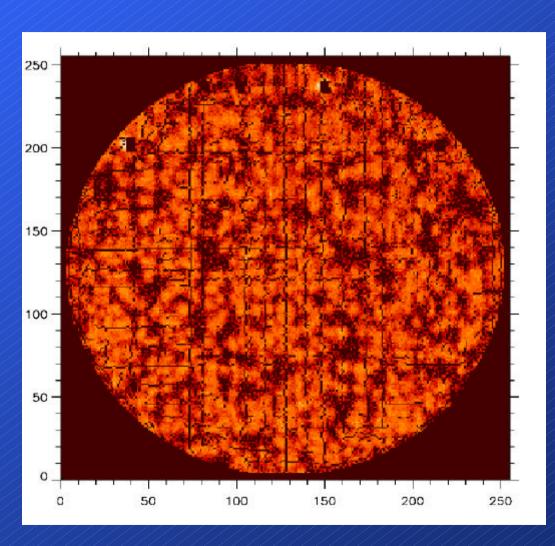
```
jmx1_evts_shd.fits
jmx1_rsti_ima.fits
jmx1_sky_ima.fits
jmx1_find_res.fits
jmx1_srcl_res.fits
```

shadowgrams
'skew' images
WCS images
individual find results
source list summary

and the stage is set for the spectral/timing analysis ...

Crab on axis in log scale

Rev 102 Log scale 3-6 keV



Noter: hvad skal findes

```
Billeder at skaffe:
Synsfelter
Standardbillede
Fine resolution
midisky = j_ima_iros
```